

Light Sensor selection for Cherenkov Telescope Array

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for the FPI WP of CTA

NDIP conference, July 4-8 2011, Lyon, France

VHE Ground-Based γ -ray Astronomy

1. Galactic Gamma-Ray Sources:

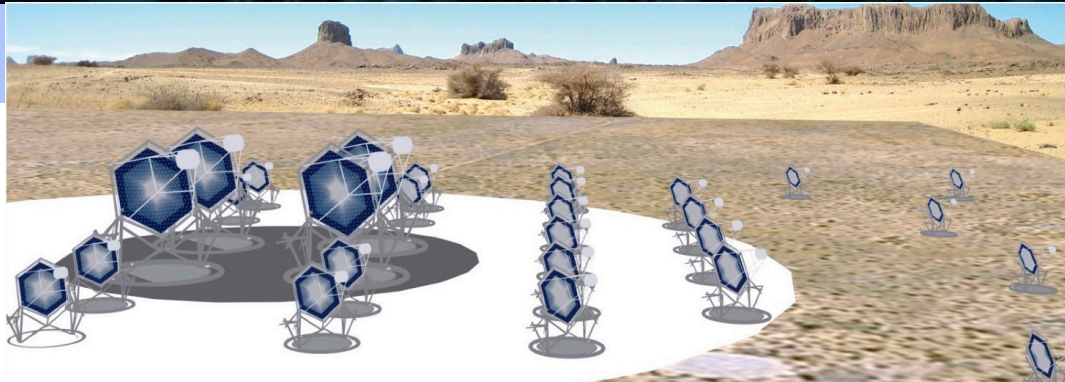
- Supernova Remnants
- Pulsars, Pulsar Wind Nebulae
- X-Ray Binaries & Micro-quasars
- Star-Formation Regions
- The Galactic Centre

2. Extragalactic Gamma-Ray Sources:

- Active Galactic Nuclei
- Extragalactic Background Light
- Gamma-Ray Bursts
- Galaxy Clusters

3. Fundamental Physics:

- Dark Matter
- Quantum Gravity
- Charged Cosmic Rays
- and more



CTA layout with different telescope sizes



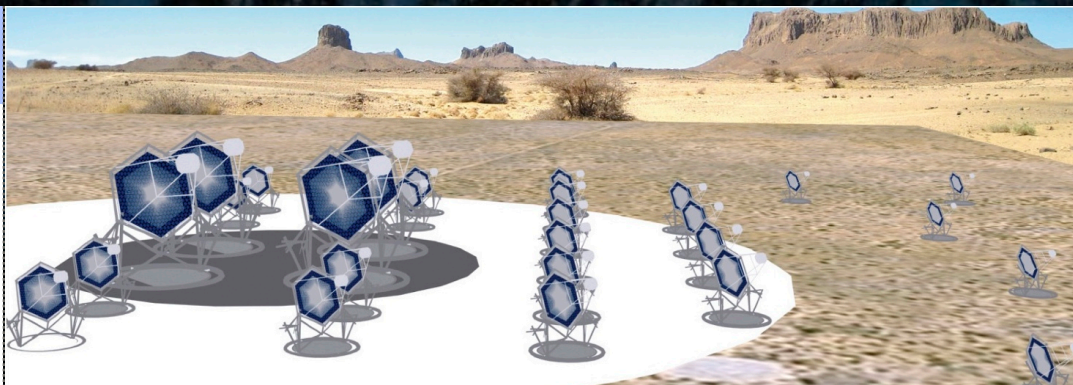
Possible CTA sites

Low-Light-Level Field

Photosensor requirements:

Total needs: $>1.5 \times 10^5$ photosensors!

- Very high Quantum Efficiency
- Almost 100 % Collection Efficiency of photo electrons from photo cathode on the 1st dynode
- After-pulsing rate: < 0.02 % above 4 Phe (needed to lower the energy threshold)
- Pulse width: $< 2 - 3$ ns
- Transit Time Spread: ≤ 1.3 ns (single photo electron)
- Stabilize photo cathode to 1st dynode voltage to ~ 350 V
- Linear Dynamic Range > 5000 photo electrons
- About 2 years ago MPI (Max Planck Institute) started a PMT development program with Hamamatsu (Japan) and Electron Tubes Enterprises (England)

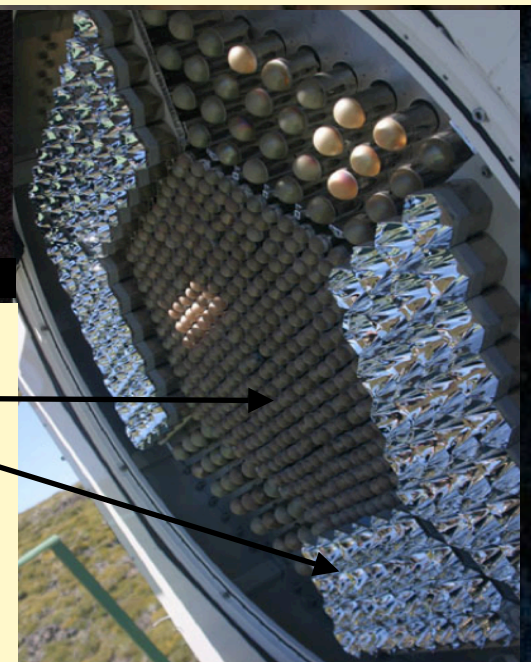


CTA layout with different telescope sizes



Bi-alkali PMTs

**MAGIC I camera as example:
Photosensors
and Light Collectors
(Winstone cones)**



FPI Main Parameter List
Summary by R. Mirzoyan
Original version: 16 April 2010;
Next modification: 27 May 2011
Latest modification: 22 June 2011



FPI WP

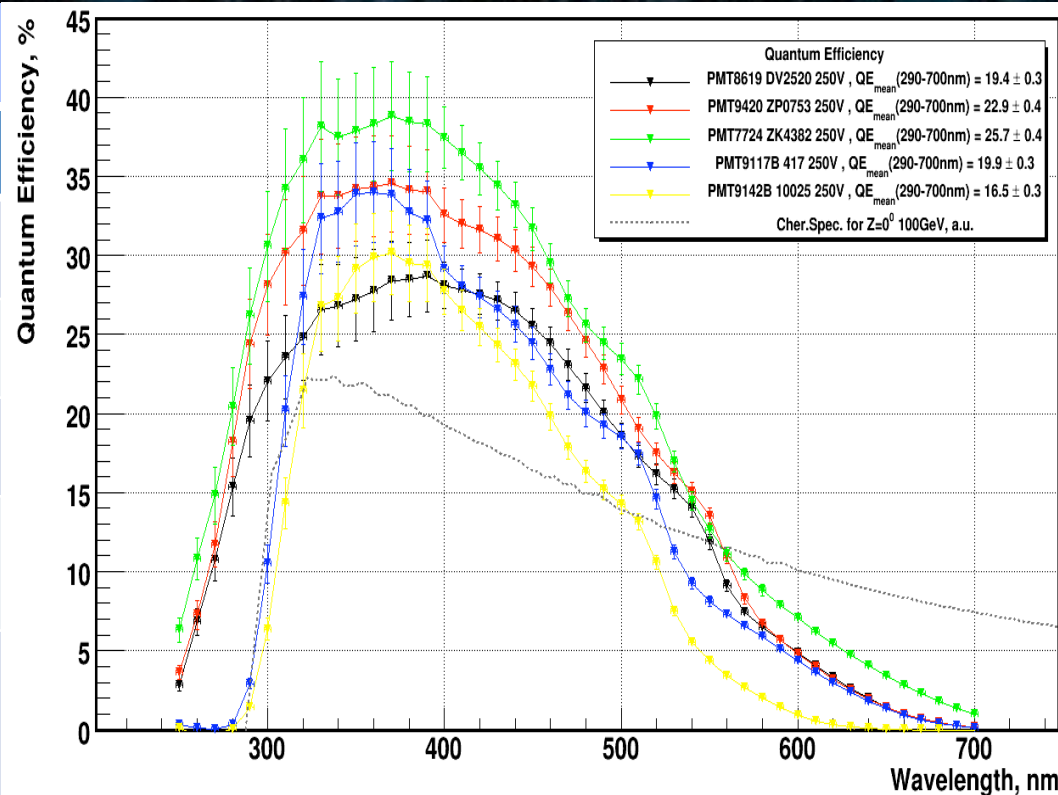
Below is shown the wish list of the main parameters of the FPI work package:

PMT:

- | | |
|--|--|
| 1. size: | 1.5 inch |
| 2. photo-cathode type: | super-bialkali |
| 3. sensitivity range: | 290 – 600 nm |
| 4. peak QE on average: | 35 % |
| 5. average QE over Cherenkov spectrum | ≥ 21 % (290-600) (without ph.e. CE) |
| 6. variation of QE over entrance window | < 10 % |
| 7. 1 st dynode ph.e. collection efficiency: | ≥ 95 % (400 nm, 30 mm diameter) |
| 8. photo cathode to 1 st dynode HV: | ≥ 300 V stabilized |
| 9. 1 st dynode gain: | ≥ 6 |
| 10. number of dynodes: | 8 |
| 11. operating gain: | $\geq 4 \times 10^4$ |
| 12. afterpulsing for threshold ≥ 4 ph.e.: | ≤ 0.02 % |
| 13. pulse width, FWHM (single ph.e.): | ≤ 2.5 ns |
| 14. transit time spread, single ph.e., rms: | 1.0 – 1.3 ns acceptable (30 mm diameter) |
| 15. single ph.e. peak/valley ratio: | ≥ 1.8 |
| 16. protected from geomagnetic field: | μ -metal shield |
| 17. dynamic range: | ≥ 5000 ph.e. |
| 18. linearity range: | ≥ 5000 ph.e. |
| 19. differential non-linearity | < 1 % |

PMT candidates

Company	Type	QE _{peak} (λ)	<QE> _{Cher}
Hamamatsu	8619 1"	28,7 ± 2,2 (390 nm)	19,4 ± 0,3
Hamamatsu	9420 1,5"	34,6 ± 3,1 (370 nm)	22,9 ± 0,4
Hamamatsu	7724 2"	38,9 ± 3,3 (370 nm)	25,7 ± 0,4
Electron Tubes	9117B 1,5"	34,0 ± 3,2 (360 nm)	19,9 ± 0,3
Electron Tubes	9142B 1,125"	30,2 ± 2,7 (370 nm)	16,5 ± 0,3



Producers claim collection efficiency >90%
(direct PDE measurements still needed)

Quantum Efficiency curves.
Dashed curve: Cherenkov radiation spectrum
from 50 GeV γ -shower.

1.5 inch PMTs has been chosen as the target PMT for all 3 types of telescopes.

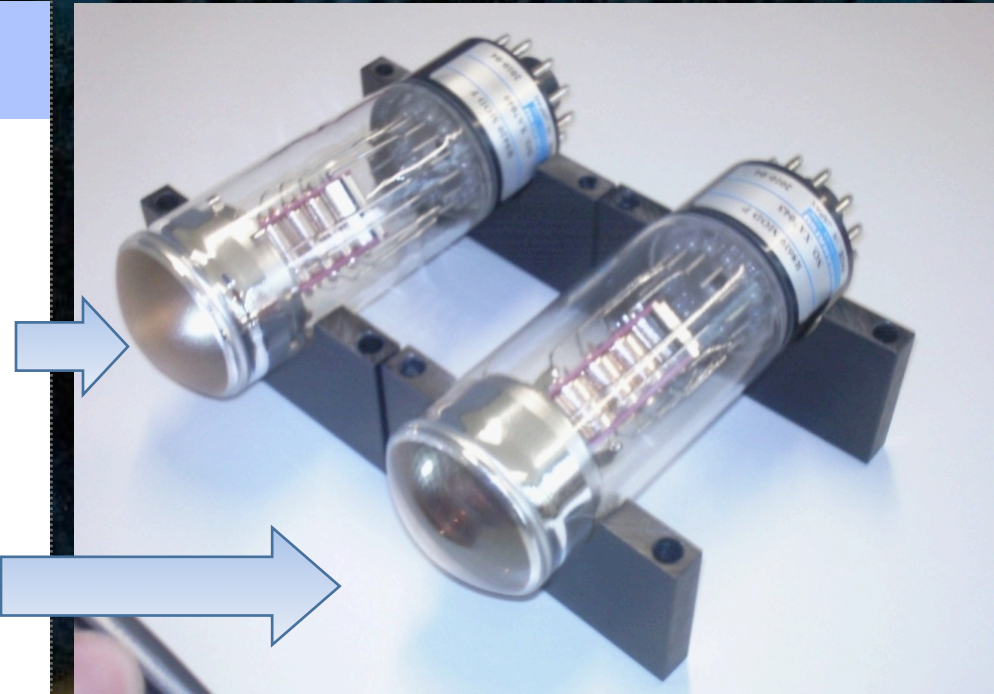
Super-bialkali (SBA) 1.5' PMTs

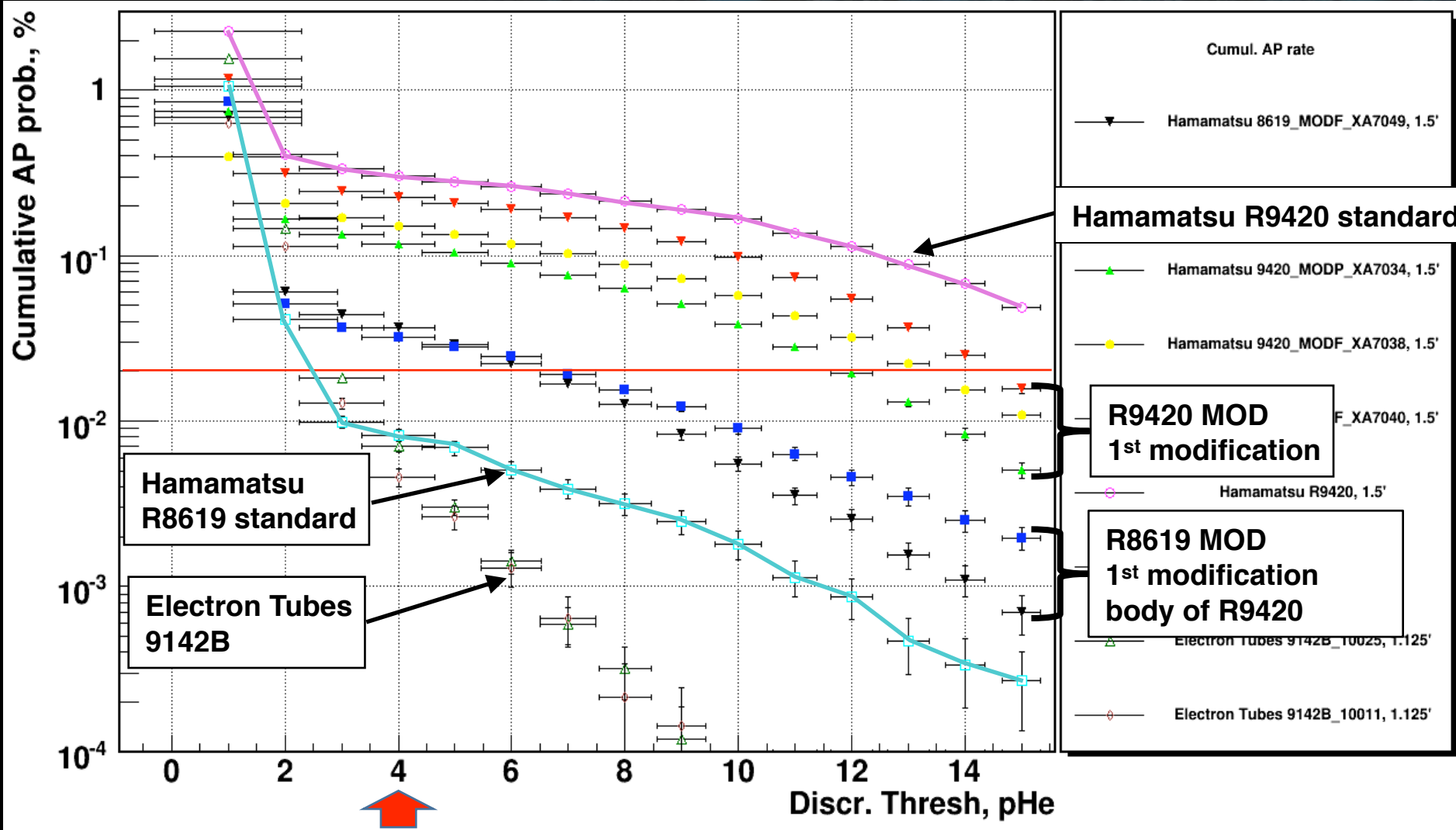
PMT samples within MPI-Hamamatsu-ET development program :

Hamamatsu R9420 modified
(convex input window shape)
(mat and polished input window types)

Hamamatsu R8619 modified
Body of R9420 but used 8619 dynode
system (also two window types)

Electron Tubes 9117B



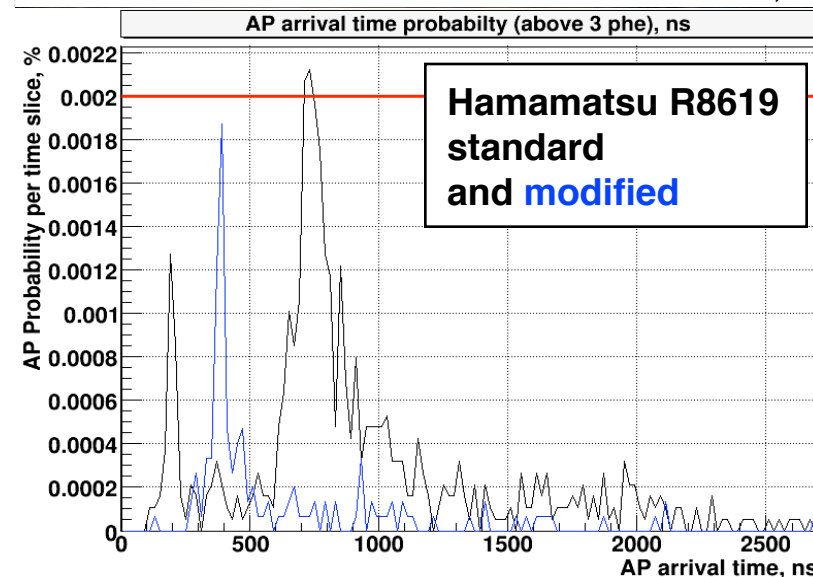
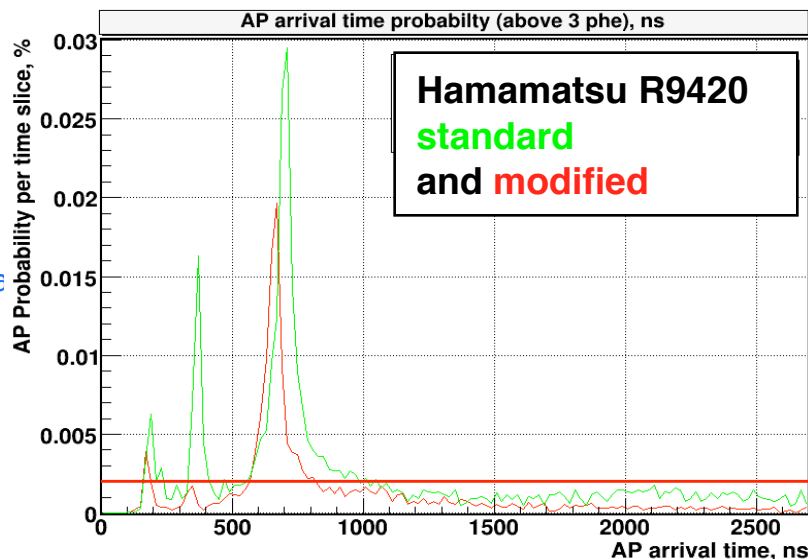
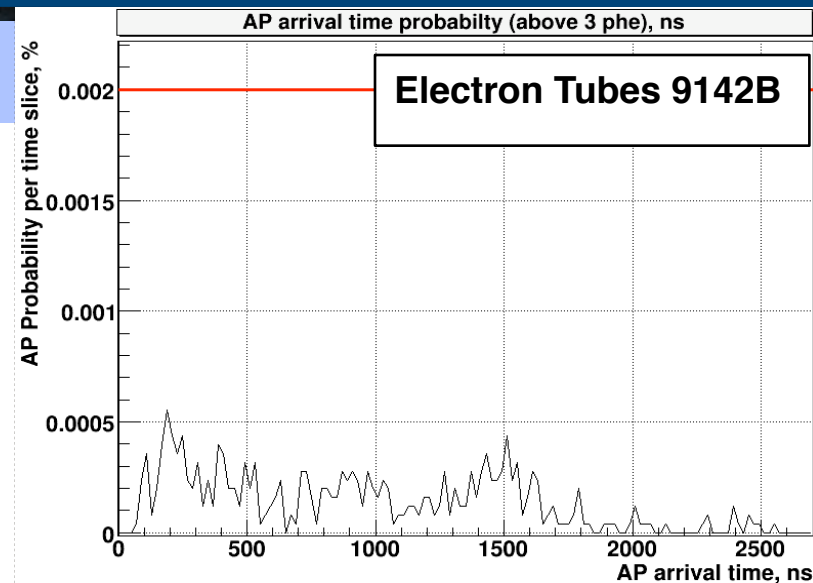


Requirement for CTA: For the threshold of 4 phe $P_{AP}(>4phe) \sim 0.02\%$ (red line)

After-Pulsing sources

Sources vary for different PMTs !

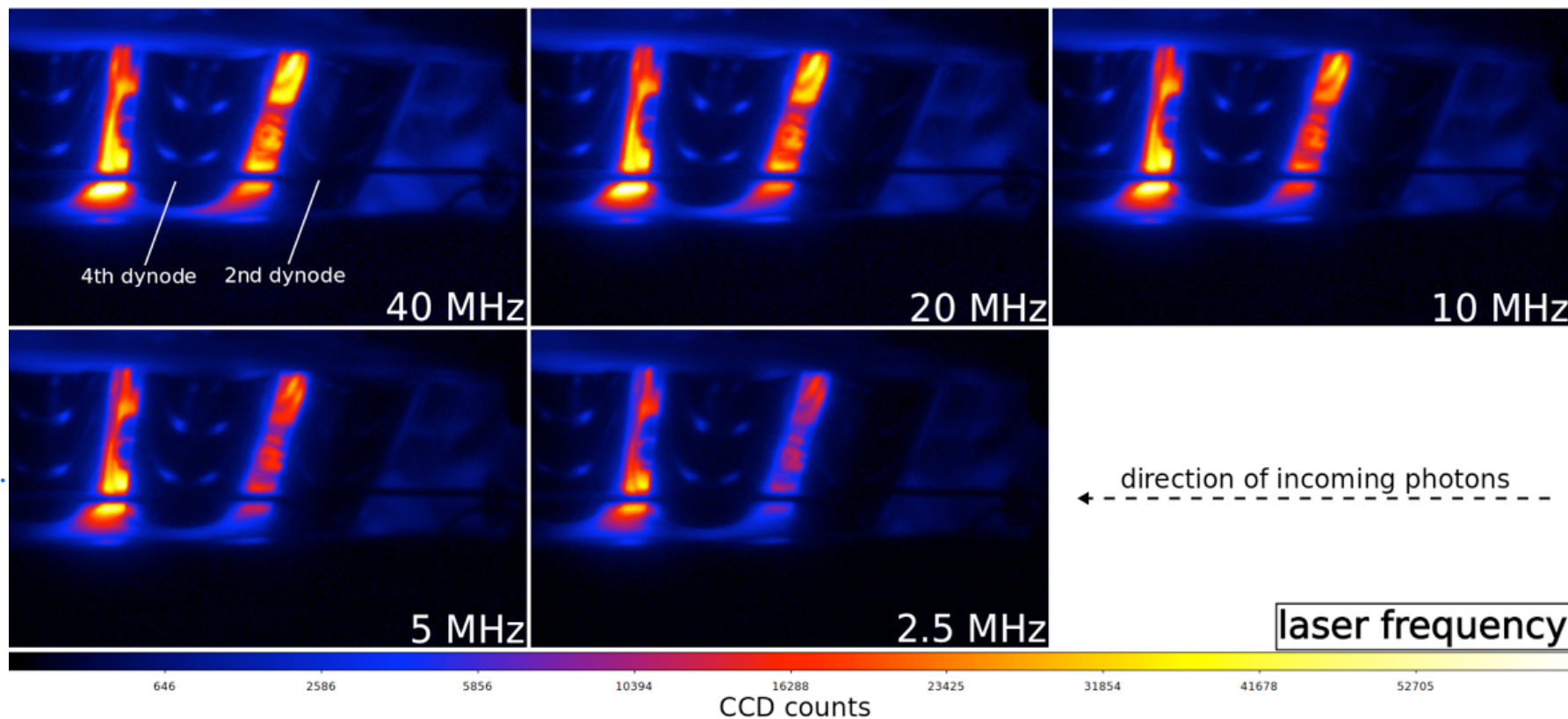
- AP arrival time measurements to distinguish sources
- Peaks: ions from first dynode (H^+, He^{+2+}, CH_4^+)
- Plateau caused by rest gases (low plateau = "clean" vacuum)
- ET9142B has very low AP rate with no pronounced peaks



Dynode system glowing

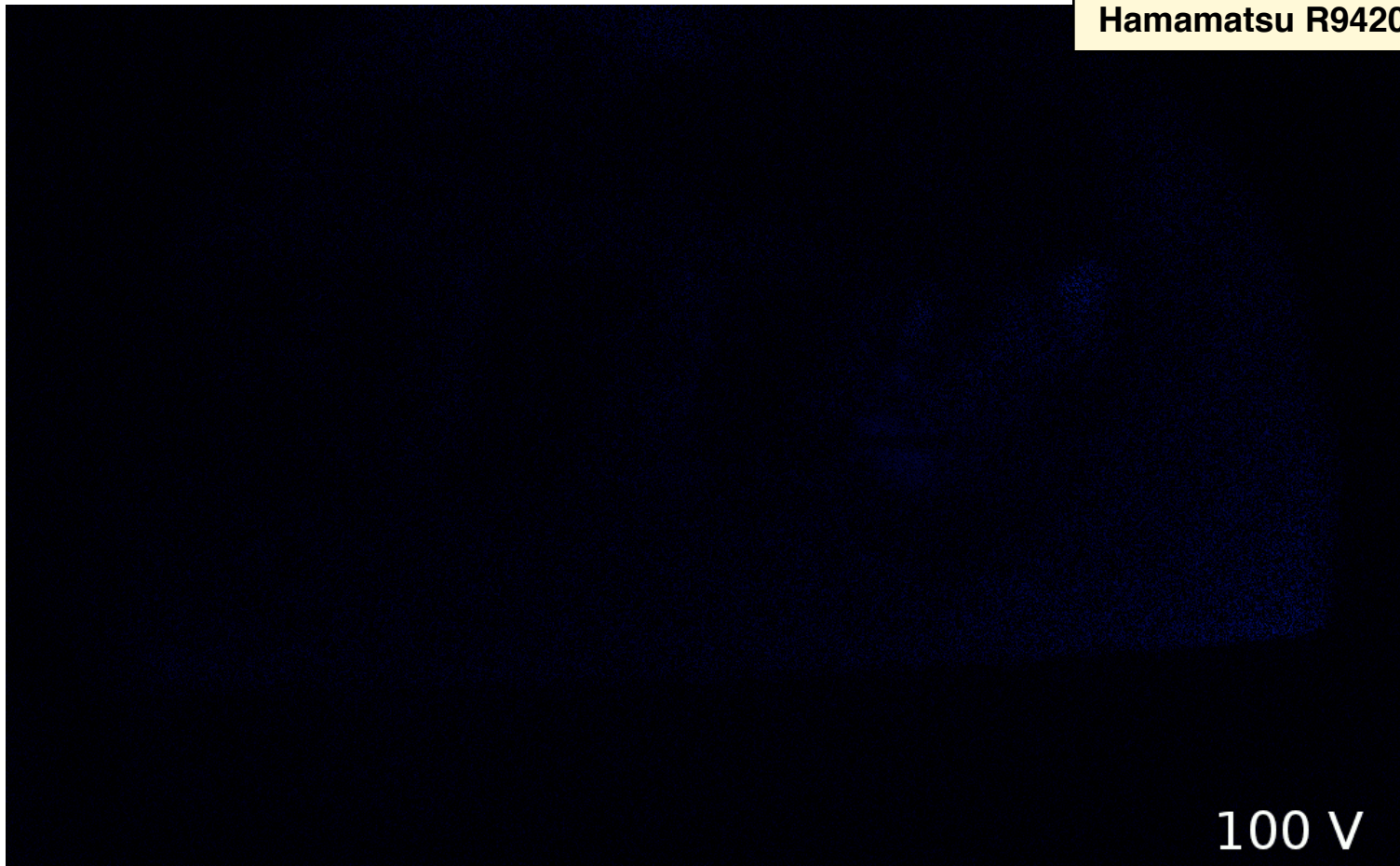
PMT irradiated by fast laser with various pulse repetition rate:

Hamamatsu R9420MOD
Exposure for every picture: 1 s



Dynode system glowing: Varying of applied voltage

Hamamatsu R9420MOD



644

2578

5838

10361

16237

23351

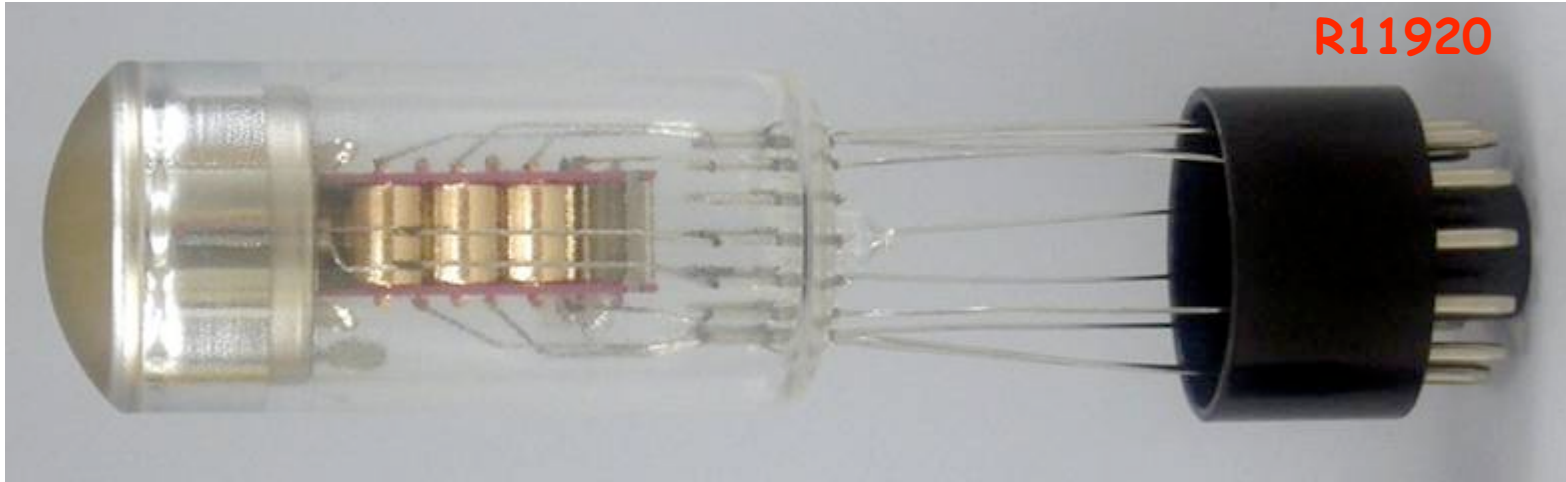
31753

41546

52539

Developed CTA PMT candidate from Hamamatsu

New name: R11920-100-01



TECHNICAL INFORMATION

TENTATIVE
Feb. 2011

R11920-100-01

For Gamma-ray Telescope (CTA project), Fast time response, CC window
38mm (1.5 inch) Diameter, Super Bialkali Photocathode, 8-stage, Head-On Type

GENERAL

Parameters		Ratings	Units
Spectral Response		300 to 650	nm
Wavelength of Maximum Response		350	nm
Window	Material	Borosilicate glass	-
	Shape	Concave-Convex (R26)	-
Photocathode	Material	Super Bialkali	-
	Minimum Effective Area	30	mm dia.
Dynode Structure / Number of Stages		Linear Focused / 8	-
Base		JEDEC No.B12-43 / Flying Lead Type	-
Operating Ambient Temperature		-30 to +50	°C
Storage Temperature		-80 to +50	°C
Suitable Socket		E678-12A (option)	-

MAXIMUM RATINGS (Absolute Maximum Values)

Parameter	Maximum Ratings	Units
Supply Voltage	Anode to Cathode	1500
	Cathode to 1 st Dynode	300
Average Anode Current	0.1	mA

CHARACTERISTICS (at 25 °C)

Parameters	Min.	Typ.	Max.	Unit
Cathode Sensitivity : Luminous (2856 K)	-	100	-	μA/lm
Cathode Blue Sensitivity Index (Cs 5-58)	-	13.5	-	-
Quantum Efficiency	at peak wavelength	32	35	-
	from 300nm to 450nm	25	-	-
Collection Efficiency (at 400nm)**	-	92	-	%
1 st Dynode Gain	6	10	-	-
Anode Sensitivity : Luminous (2856 K)	-	4	-	A/mm
Gain	-	4x10 ⁴	-	-
Single Photon counting Peak to Valley Ratio	1.8	2.5	-	-
Anode Dark Current (after 30min storage in the darkness)	-	5	20	nA
After Pulsing (threshold 4p.e. and Gain 4x10 ⁴ voltage)	-	0.05	-	%
Anode Pulse Rise Time**	-	2.3	-	ns
Anode Pulse Width (FWHM)**	-	-	3.0	ns
Electron Transit Time**	-	23	-	ns
Transit Time Spread (FWHM with single p.e.)**	-	-	1.3	ns
Pulse Linearity (+/-2% deviation)	15	20	-	mA
Life (50% drop in Gain)	200	-	-	C

NOTE : Anode characteristics are measured with a voltage distribution ratio and supply voltage shown below.

(** Collection Efficiency and Time response are defined with effective area of 30 mm in diameter.)

TECHNICAL INFORMATION

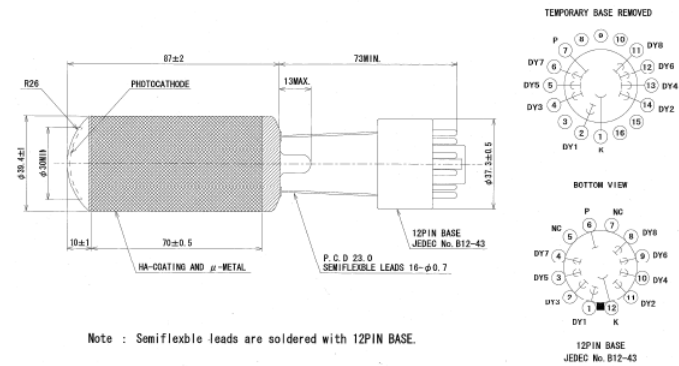
TENTATIVE
Feb. 2011

VOLTAGE DIVIDER AND SUPPLY VOLTAGE

Electrodes	K	Dy1	Dy2	Dy3	Dy4	Dy5	Dy6	Dy7	Dy8	P
Ratio	300V(zener diode)	1	2	1	1	1	1	2	1	

Supply Voltage: 1000 V, K: Cathode, Dy: Dynode, P: Anode

DIMENSIONAL OUTLINE AND BASING DIAGRAM



Note : Semiflexible leads are soldered with 12PIN BASE.

Figure 1: Dimensional Outline (Unit: mm)

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No warranty, expressed or implied, is created by furnishing this information.

These parameters are under improvements

25mm->20mm convex window. CE improves, TTS slightly degrades.

Please follow industrial wastes in

Additional material may be reliable. However, please without notice

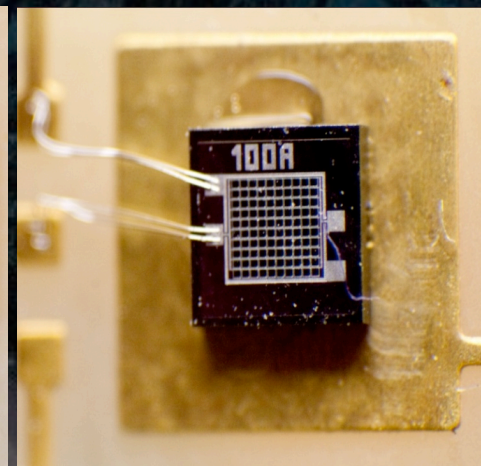
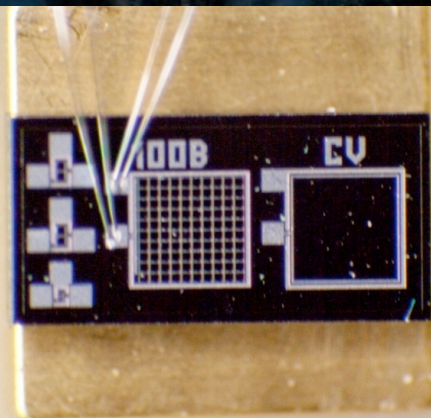
First SiPMs candidates

MEPhi - MPI samples:

- Different pitch (50μ , 100μ), p on n type
- Two SiPMs sizes: (1 & 3 mm²)
- Operation voltages 26V and 35V
- In total ~20 different types of SiPMs

Also tested:

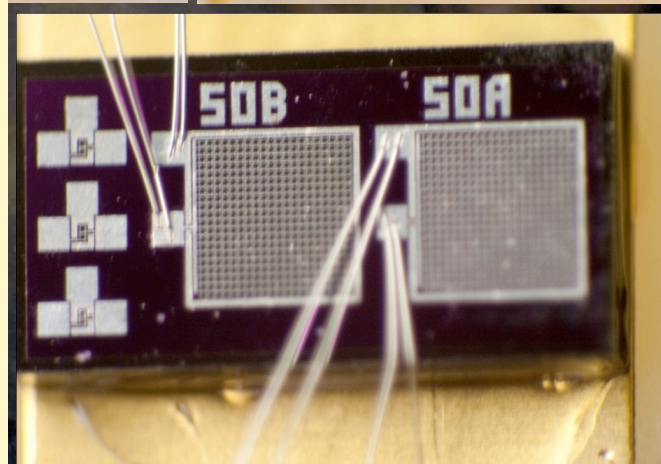
- Hamamatsu MPPCs, 1 & 3mm, S10362-11-50U
- Planning to test other types when available e.g. Philips, and other manufactures



First prototypes from MEPHI:

1x1mm, 35V and 26V bias voltages.
100 μ and 50 μ pitch.
No trenches

B, A = different geometrical efficiencies
(B-high, A-lower)

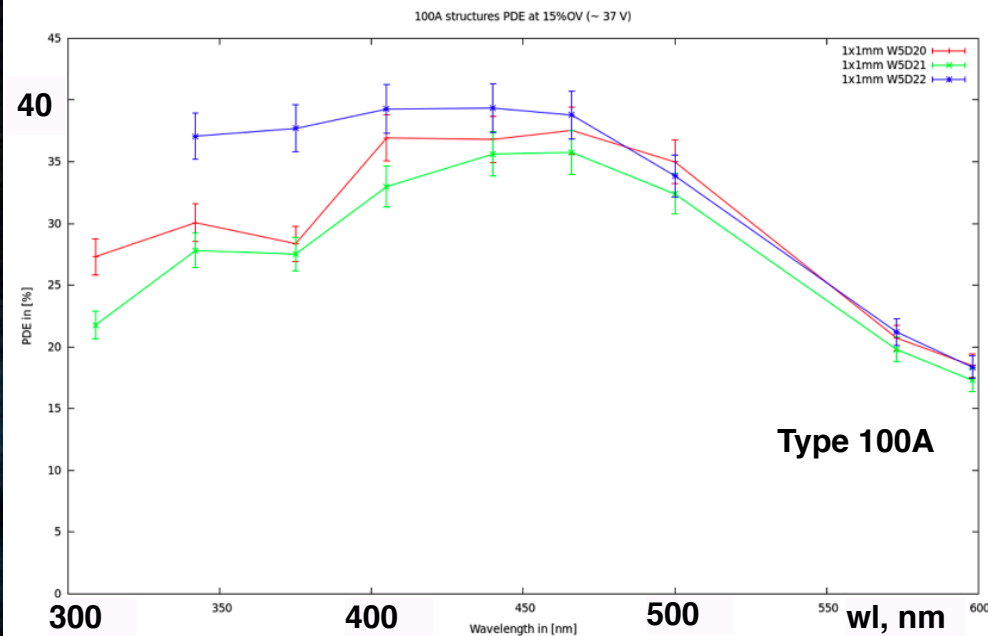


Updated SiPMs candidates

Updated MEPhi - MPI (EXCELITAS) samples:

- Type 100A and 100B
- Improved PDE in UV region
- X-Talk suppression: Trenches, second p-n junction, ion implantation...
- Measurements of type A reveal samples with high (and pretty “flat”) PDE:

PDE(400nm)~40% !!!



Measurement of PDE for applied over-voltage $(V-V_b)/V_b=15\%$

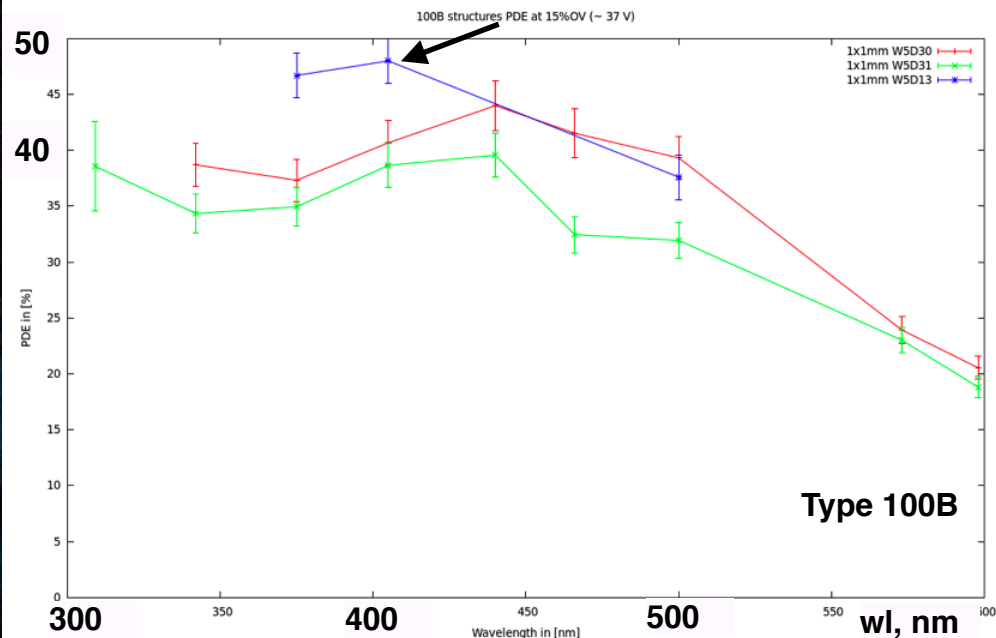
Updated prototypes Type 100A ($U_b=37V$):
Improved UV sensitivity. 1x1mm,

Updated SiPMs candidates

Updated MEPHI - MPI (EXCELITAS) samples:

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- X-Talk suppression: Trenches, second p-n junction, ion implantation...
- Measurements of type B reveal samples with high (and pretty “flat”) PDE:

PDE(400nm)~45% !!!



Measurement of PDE for applied over-voltage $(U-U_b)/U_b=15\%$

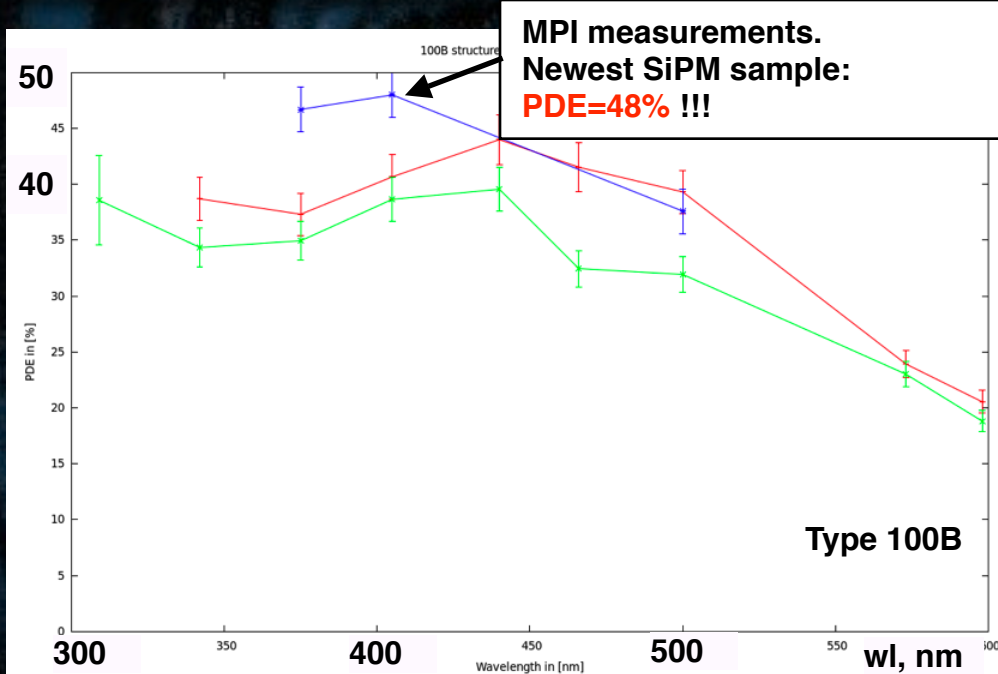
Updated prototypes Type 100B ($U_b=37V$):
Improved UV sensitivity. 1x1 mm,

Updated SiPMs candidates

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Measurement of PDE for applied over-voltage $(U-U_b)/U_b=15\%$

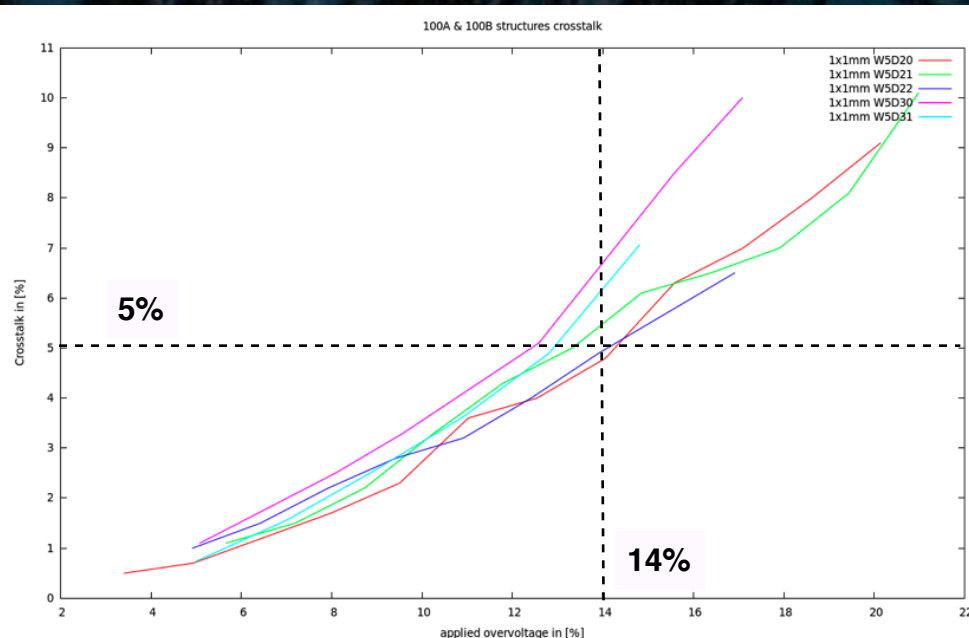
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Updated MEPhi - MPI (EXCELITAS) samples:

- Type 100A and 100B
- Improved PDE in UV region
- X-Talk suppression: Trenches, second p-n junction, ion implantation...
- Cross-talk measurements: Very low crosstalk of

$P_{X\text{-talk}}(\Delta U/U_b=14\%) \sim 5\% !!!$



Optical cross-talk versus over-voltage $(U-U_b)/U_b$
(PDE saturates at 10-15% over-voltage)

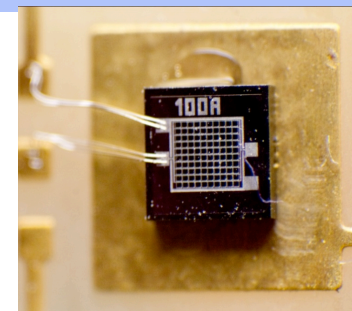
Updated prototypes Type 100B ($U_b=37V$):
Improved UV sensitivity. 1x1mm, Optical trenches
introduced

PMTs:



SiPMs:

- One may assume that it could make sense to start using SiPMs in our telescopes once their efficiency will exceed the PDE of PMTs by 1.5-2 times at comparable costs.
- SiPMs are becoming serious sensor candidates for the CTA project:
PDE ($\Delta U/U_b=0.15$) $\sim 50\%$ reached for low cross-talk $P_{X\text{-talk}}(\Delta U/U_b=0.14) \sim 5\%$!
- Intending to evaluate all new SiPM candidate sensors
- We believe in a few years one can have high quality SiPMs from several manufacturers that can be used also in our project.



PMTs:

- PMT development for the needs of CTA is in progress
- Already now the PMTs of Hamamatsu came rather close to the specified target parameter values: peak PDE $\sim 32\%$, After-pulsing slightly above 0.02% .
- Next two years further optimization of PMTs from both Hamamatsu and Electron Tubes Enterprises (Some financial support from CTA is foreseen)
- Design and tests of PMT-based pixels and clusters of 7-pixels



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