

# An ebCMOS camera system for marine bioluminescence observation: the « LuSeapher » prototype.

## IPNL ebCMOS group

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## CPPM Antares/Biocam

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Detector R&D in collaboration with IPHC Strasbourg and PHOTONIS

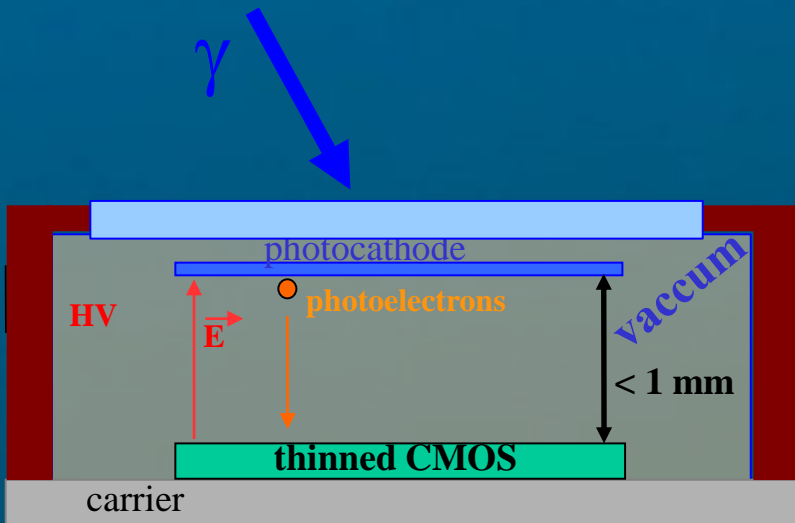


# General ebCMOS Camera System

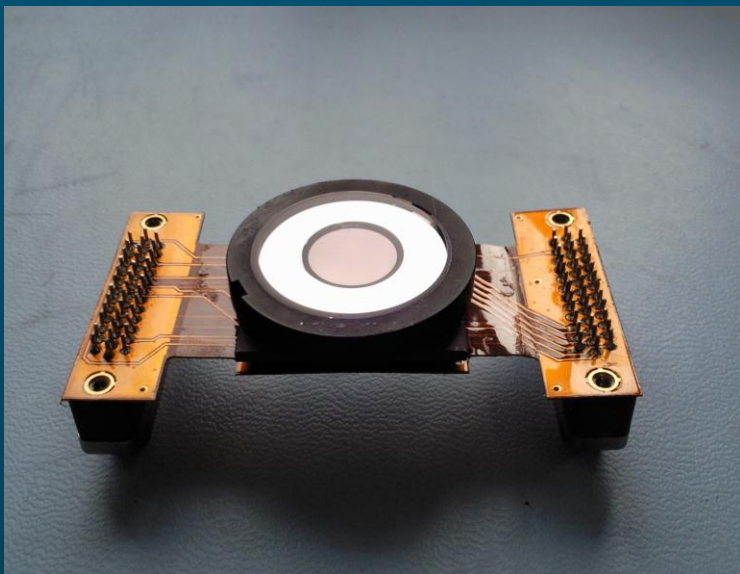
- ebCMOS : *electron bombarded Active Pixel Sensor CMOS*
  - Photoconversion: Photocathode vacuum tube proximity focusing
  - Multiplication gain: accelerating electric field
  - Readout: back thinned APS CMOS array
- Detection of a **low light** source in dark environment with ultra low noise detector
  - Dark Count: less than  $10^{-4}$  photon/pix/frame.
  - Single-Photon sensitivity and photon counting capability
- High readout speed: pixel clock = 2.5-40 MHz, Frame rate = 62-500 Hz
- « LuSeapher » is the final camera system



# ebCMOS working principle



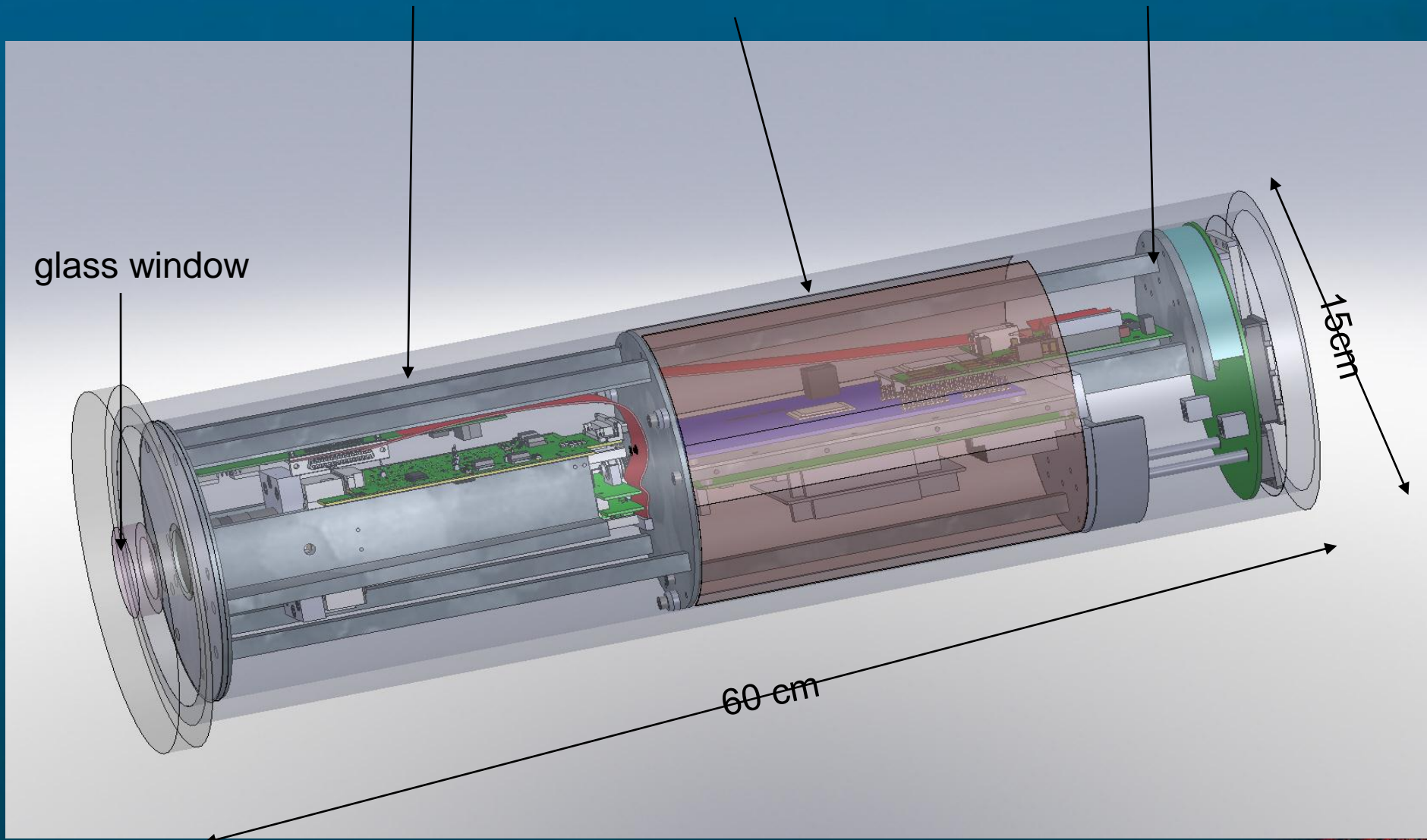
- 400x400 pixels chip (0.25 micron)
  - 10  $\mu\text{m}$  pixel pitch, 3T architecture
  - detection size : 16  $\text{mm}^2$
  - 80 nm dead layer (1.2 keV), 8  $\mu\text{m}$  sensitive layer
- gain with HV 2.8 kV
  - SinglePhotoelectron Gain:  
(2.8-1.2)keV/3.6 eV=444  $e^-$
  - Charge Collection efficiency = 60%
  - Coll. Time  $\approx$  100 ns
  - Fake photon due to CMOS noise:  
 **$10^{-5}$**
- S20 (Multialkali) cathode
  - 25% @ 450 nm



# Smart Camera System's design

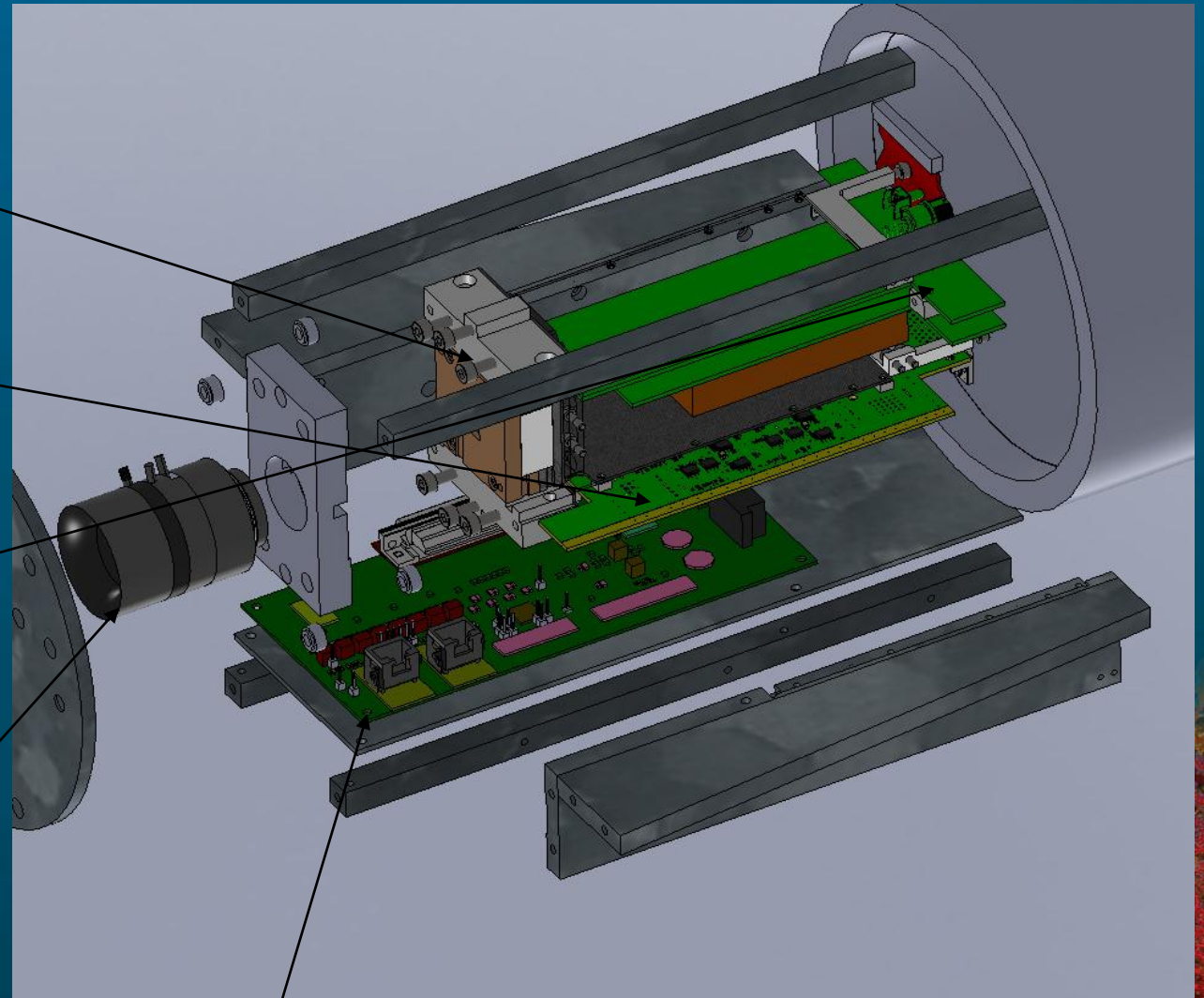
Titanium waterproof tube : 15cm diameter and 60cm long

3 main parts : lens and ebCMOS, DAQ and CPU boards and power supply



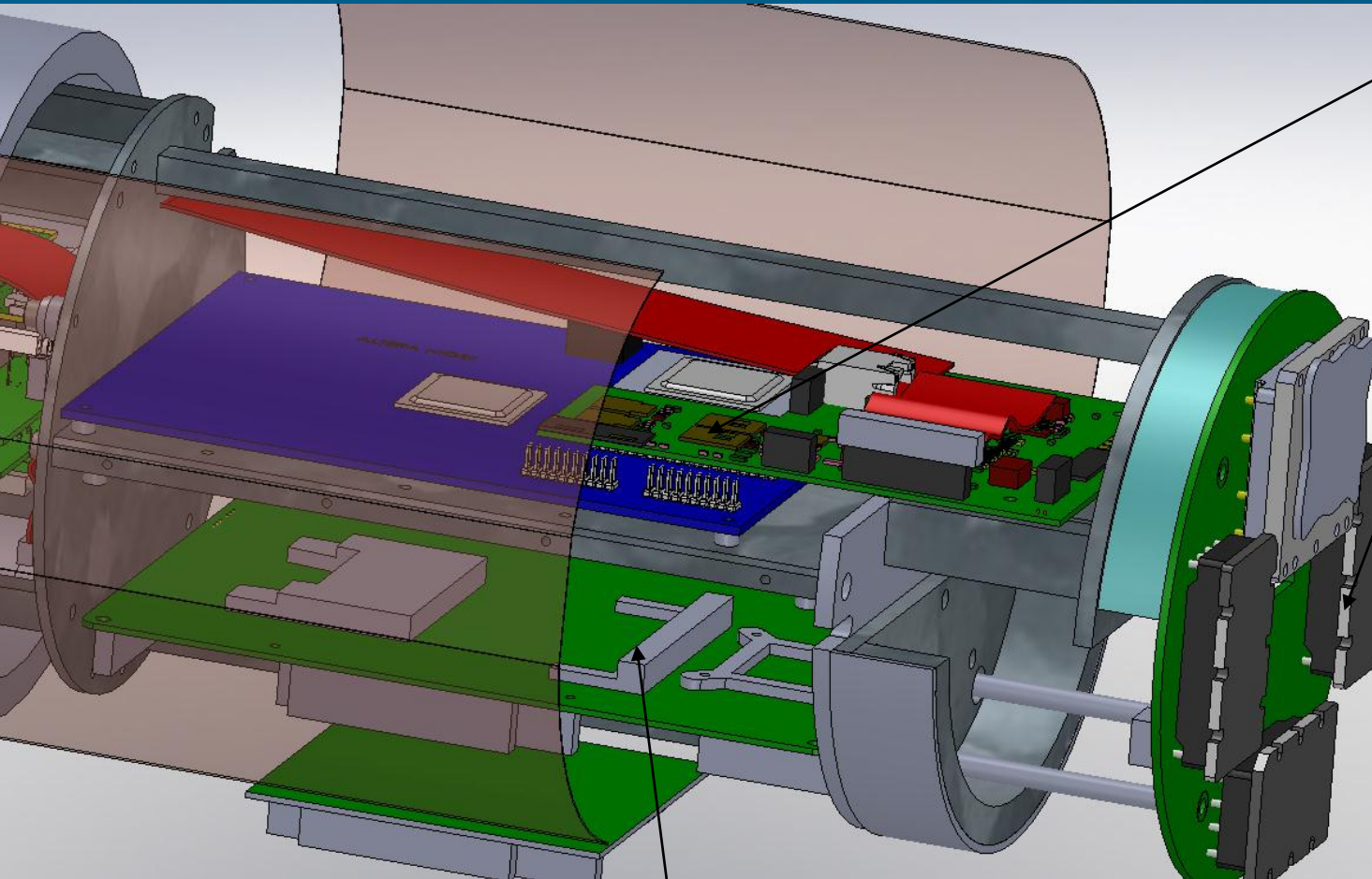
# Analog part

- ebCMOS
  - 400x400
  - 3 kV
  - S20 photocathode
  - cooled : 15° C
- sensor frontend board
- control board
  - CMOS operating voltage settings
  - chip temperature control
  - HV control
  - PS management
  - monitoring
- varifocal lens (Tamron)
  - focal 6 mm
  - F/D=1.2
  - $d_{\min} = 30$  cm



interface board to DAQ

# DAQ part



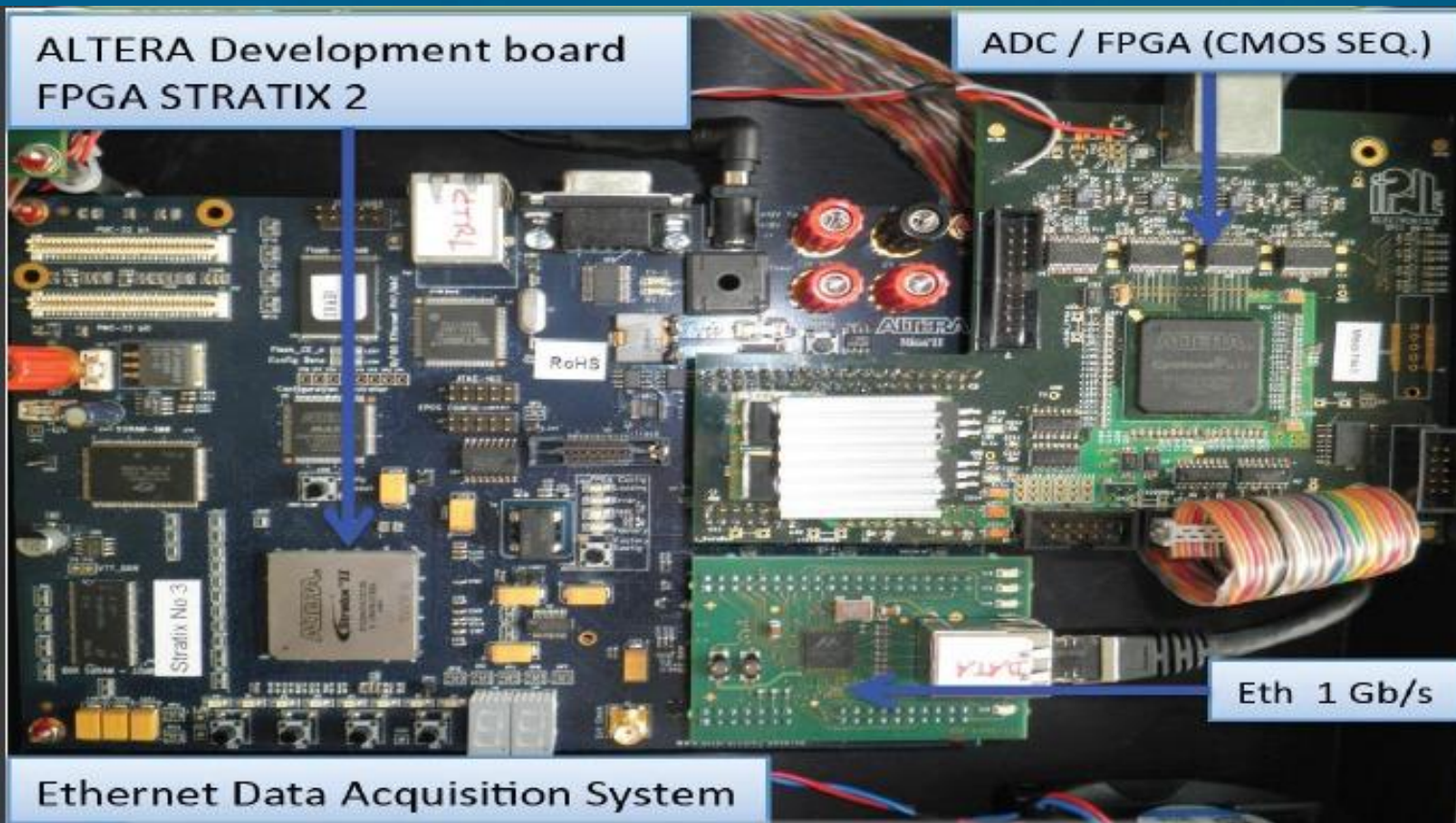
- DAQ (next slide)
- PS units
  - input: 48V
  - PC 12V
  - DAQ: 12V,9V,3.3V...
  - Slowcontrol : 12V
- power consumption : 100 W

PCM 9562 Single Board Computer

Atom D510 dual core 1.67 GHz

2 GB RAM, 32GB CF card for OS, 64 GB SSD for data

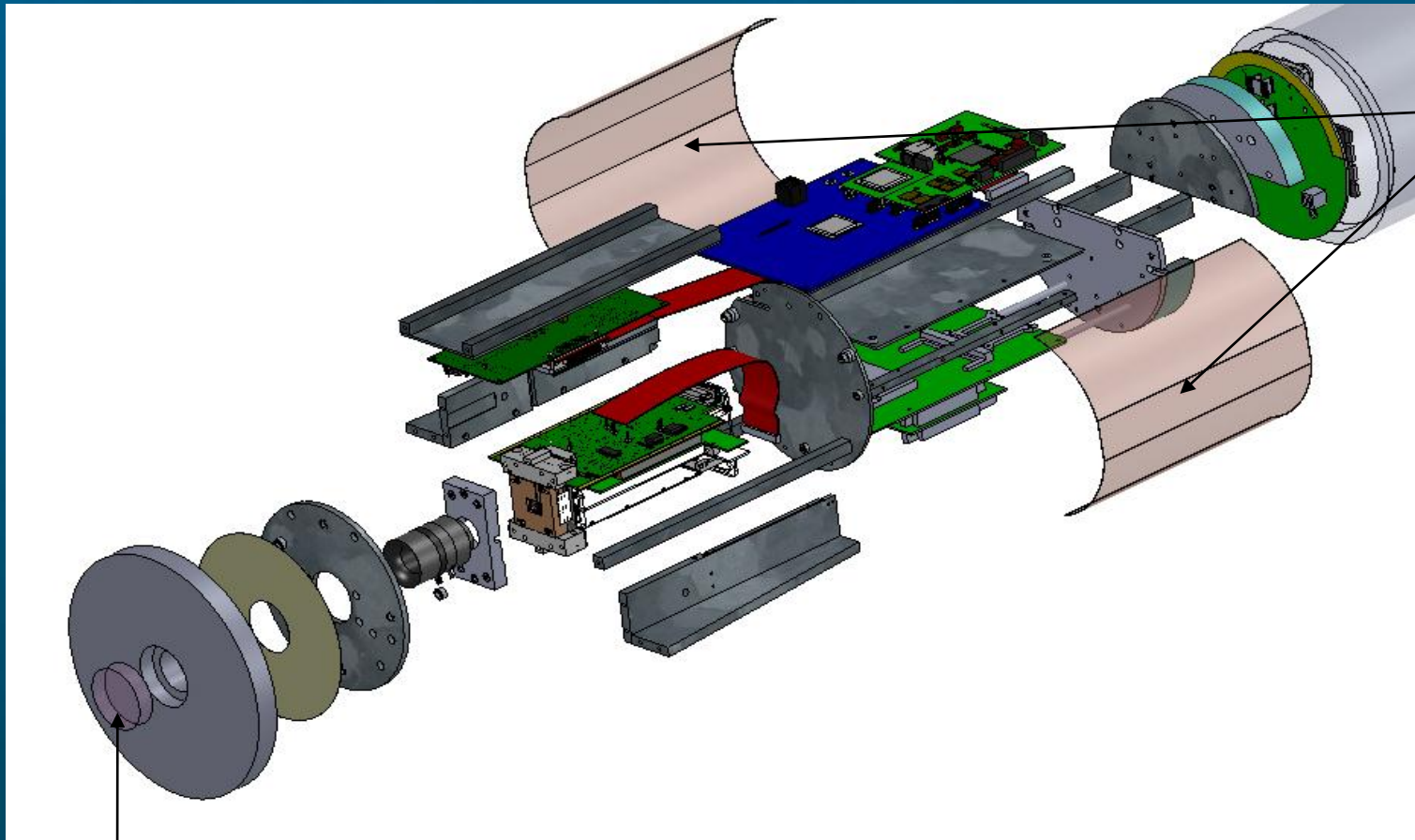
# DAQ board



low 5 MHz readout with 4 channels for 400x400 pixels gives 125 fps framerate, 2 frames needed for CDS subtraction → 62.5 fps framerate

2 mezzanines : 4 ADCs 12 bits, 1 ethernet 1Gbit/s

# Spread View



heat dissipation by contact with Titanium and 12° sea water

CEM isolation from tube

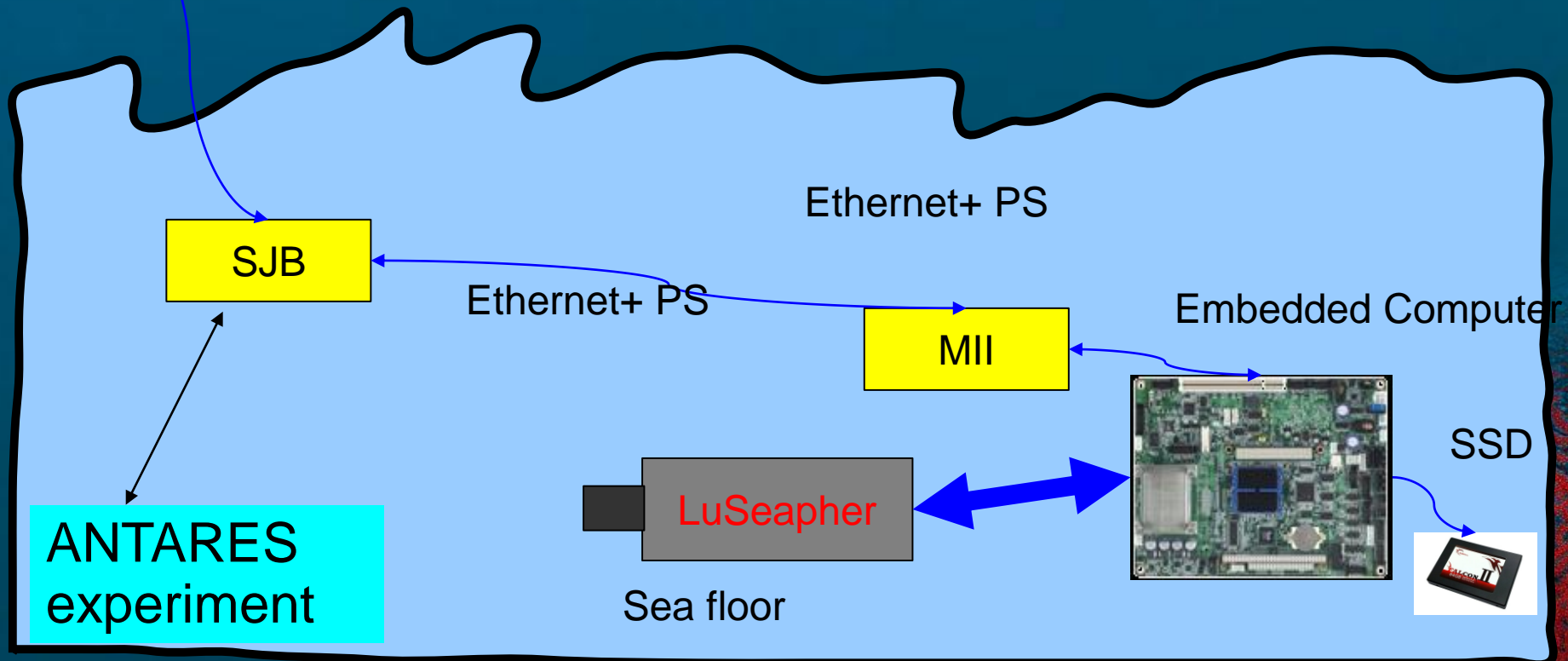
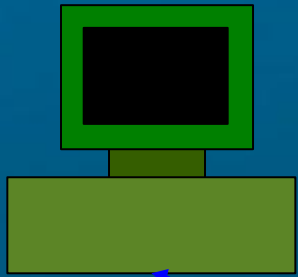
3cm diameter glass window : small FOV





# schematic view

- Antares is a cabled sea-floor observatory dedicated to neutrino physics. it's located 2500m depth off the mediterranean sea near Toulon (France).  
LuSeapher device is connected to the SJB nearby.
- MII is a multidisciplinary experiment connected to the BJS
- The ebCMOS is one of the sensor put on the MII

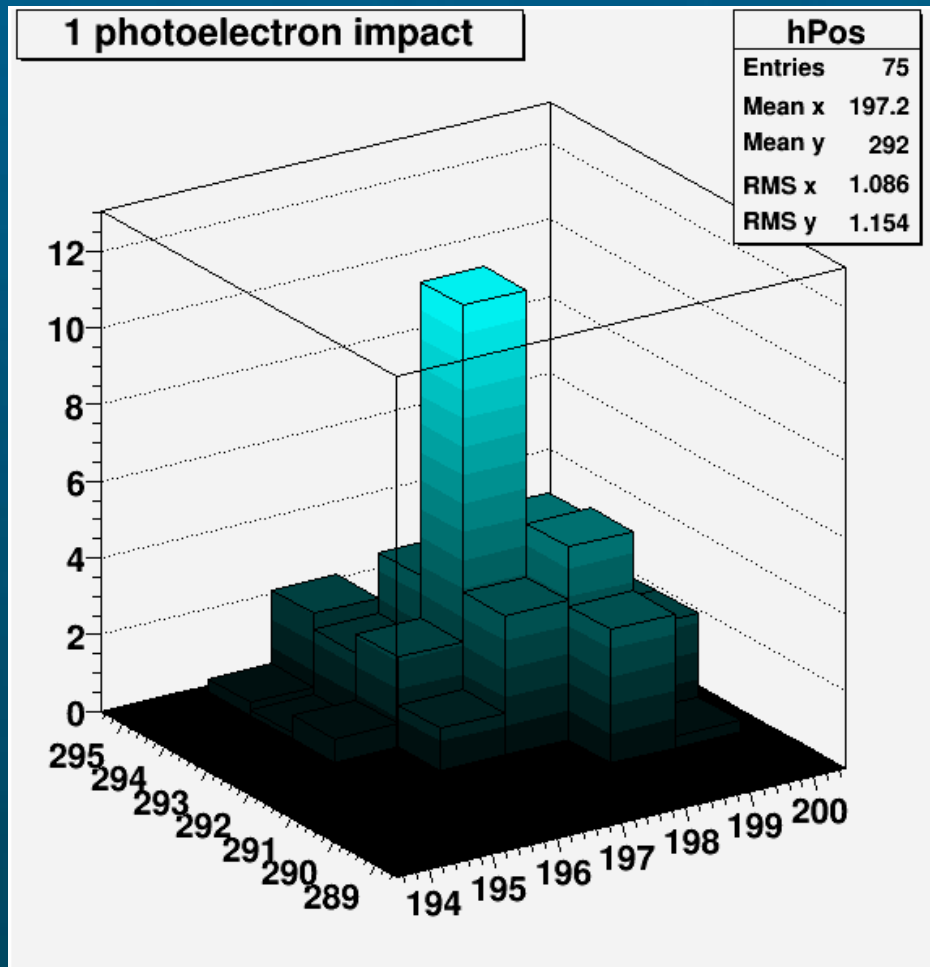


# operating mode

- bioluminescence observation in sea darkness
- due to a limited bandwidth, stream readout cannot be stored on disk (400x400 pixels @ 62.5 fps with 1 byte/pixel is equivalent to 600 MB/min or 80 Mbits/s).
- gain (HV) is ON and data are buffered on local (SSD) disk until the next trigger, then the circular buffer is read out while signal remains above the threshold.
- the trigger is very simple : no assumption on the size or the form of the signal. Data transfer if the number of reconstructed photons is above 14.
- Data storage 50 (1s) frames before trigger and 300 (5s) frames after the object comes out the field of view.
- typical duration : 1s + 1s (real event) + 5s = 7s, 70 Mbytes

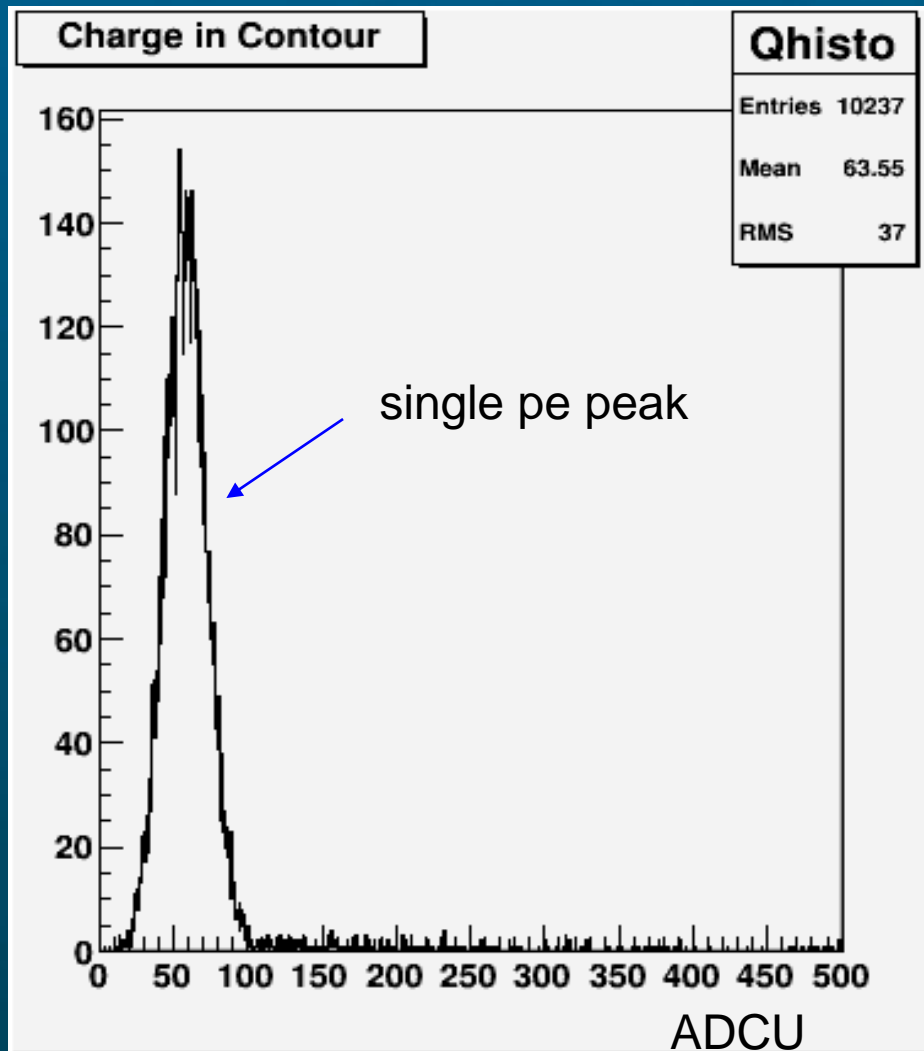


# photons reconstruction



- 5x5 pixels cluster for one photoelectron (pe)
- photoelectron pattern
  - 1 central seed  $Q > 3$  ADCU
  - 30% of pixels  $> 50\%$   $Q_{seed}$
  - 50% of pixels  $> 30\%$   $Q_{seed}$
- efficiency, purity : 96%
- CMOS temporal noise contribution to p.e. trigger rate is removed by this simple pattern recognition.
- Center of Gravity (COG) calculation gives  $2\mu\text{m}$  resolution (20% pitch)

# photoelectron noise measurement



- dark count  $\rightarrow$  0.78 pe /frame
- Ion feedback (residual ions in vacuum tube hitting the photocathode) additional signal less than  $5e-3$ /frame (0.3 Hz)
- trigger threshold : 14 pe (could be less with Ion feedback identification algo)

# Results

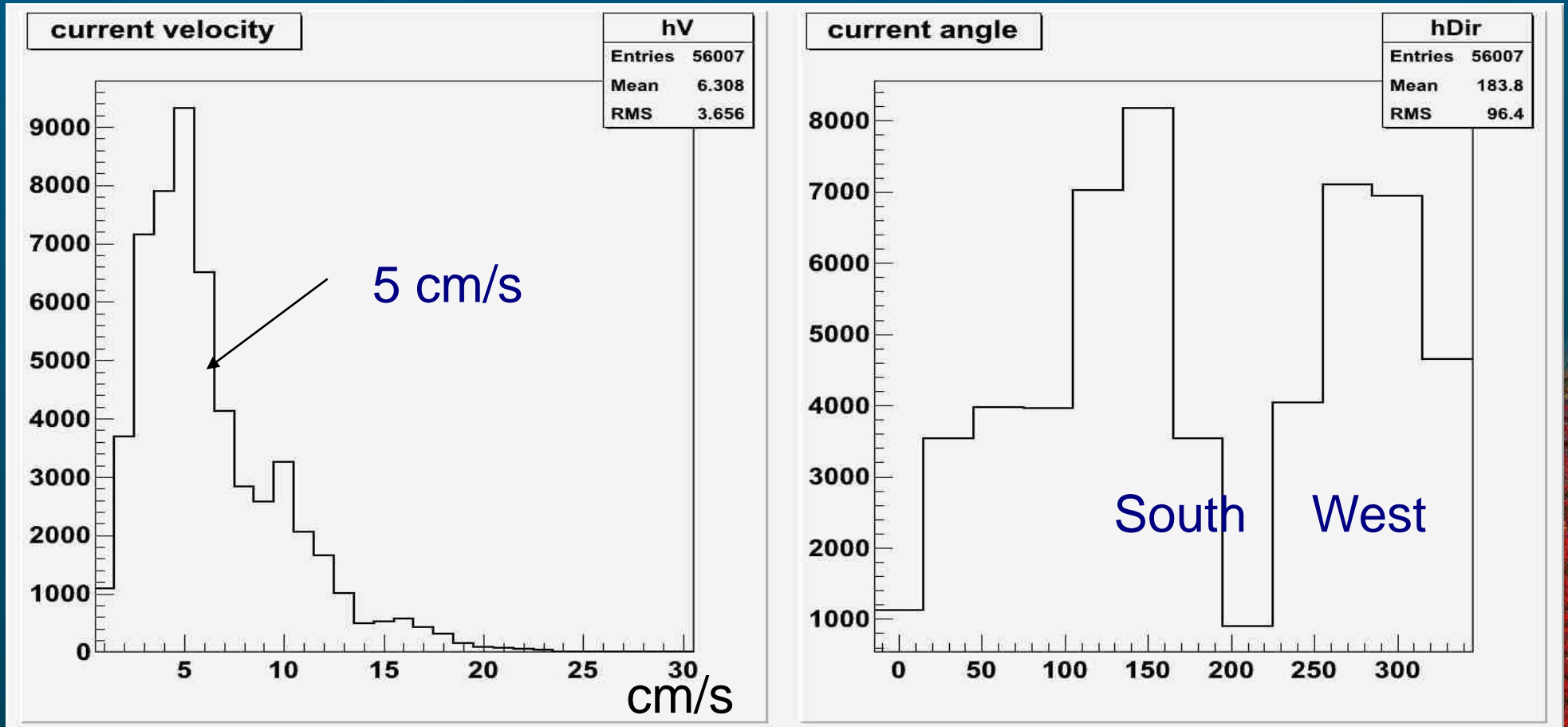
- LuSeapher is fully fonctionnal since middle of december
- photon counting imaging experiment :
  - photoelectron position (X,Y) at a function of time (it's the new contribution from this luseapher version)
  - Number of photoelectrons vs time
- about 4000 events occurs, drastic selection :
  - 65% of the total are 1 frame event → rejected
  - 15% are bad events (acquisition defaults, bad pedestals substruction, sensor's saturation)
  - 5% are not synchornized ( peak at t=70 frames)
  - 622 events (15%) are kept for the analysis

goal :

- ✓ characterization and identification of the observed patterns (size and relaxation time)
- ✓ find correlation with oceanographics conditions (sea current velocity and direction)

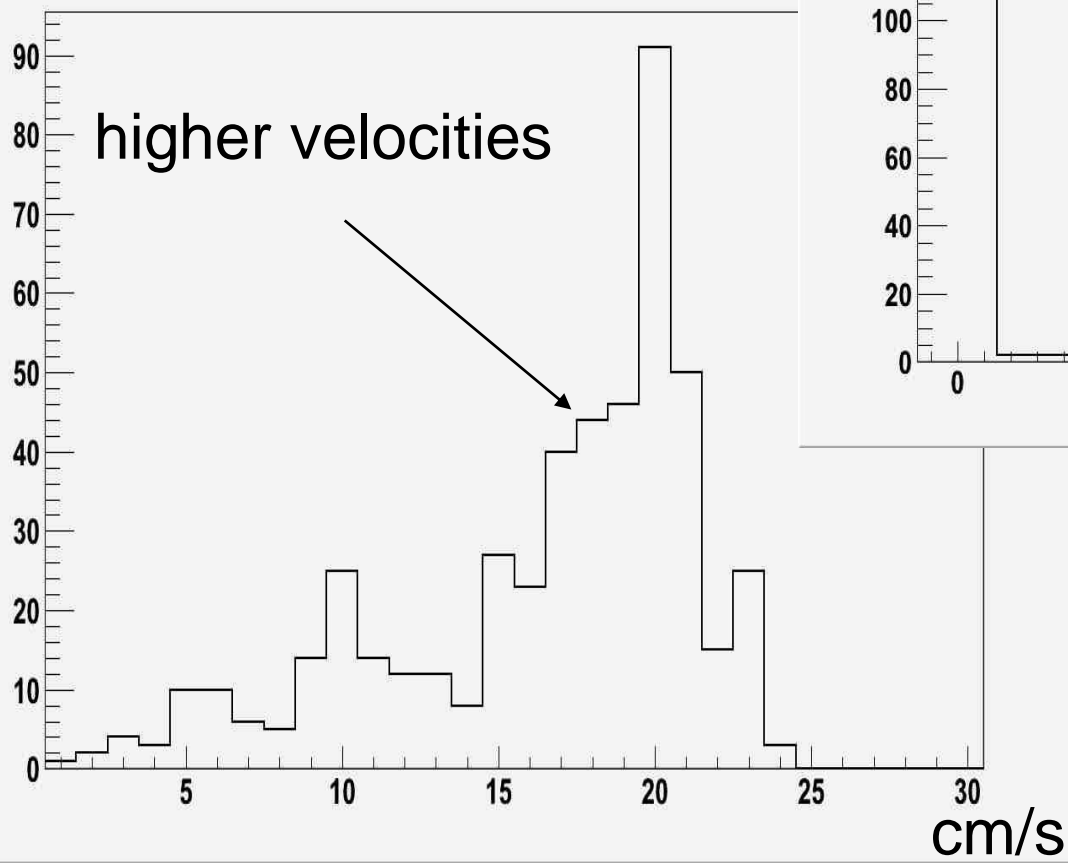
# Basic Oceanographics conditions

- From Antares experiment
  - sea current velocity (several cm/s)
  - current direction
- full set of data since 1st jan 2011.
- might be correlated with our events.

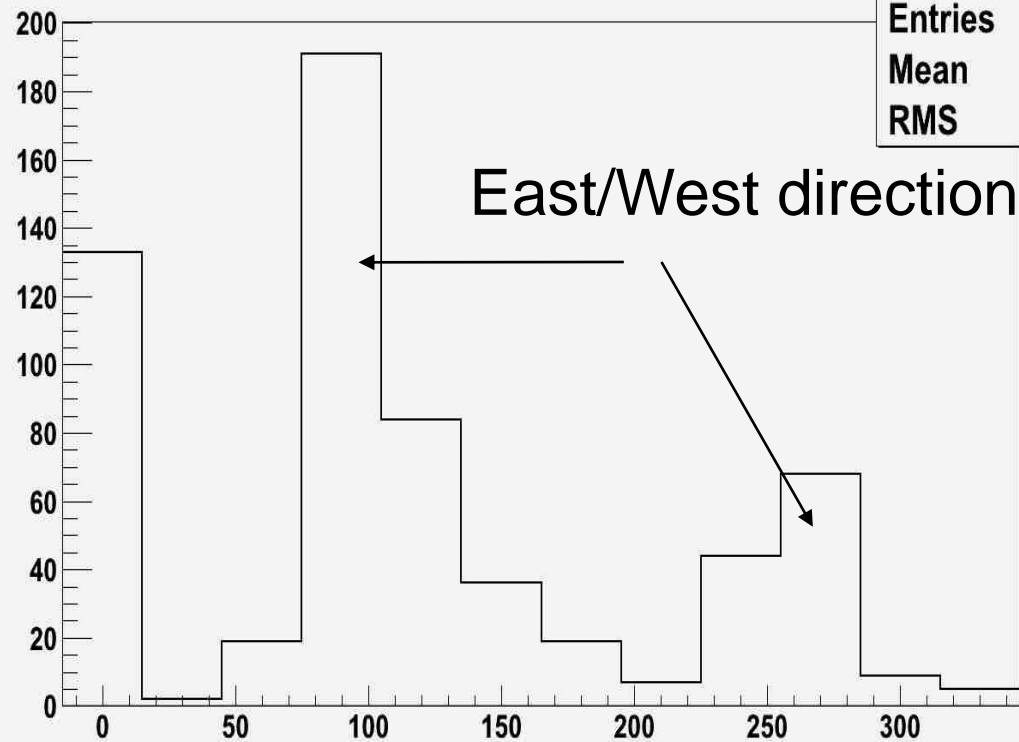


# Oceanographics conditions for LuSeapher selected events

current velocity



current angle



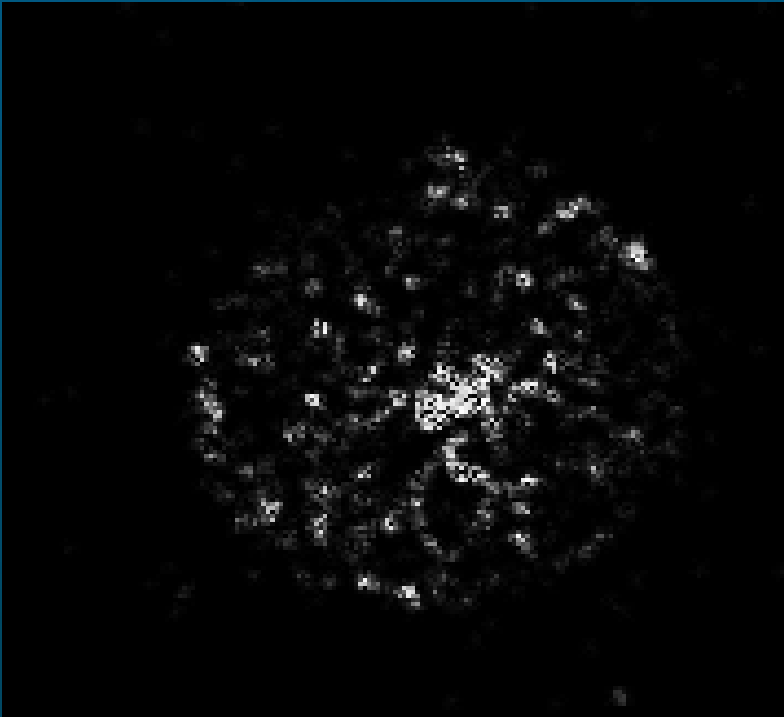
hDir

Entries	622
Mean	117.1
RMS	88.76

East/West direction



# typical event



Integrated picture of 200 frames.

more to see on videos

...

main features:

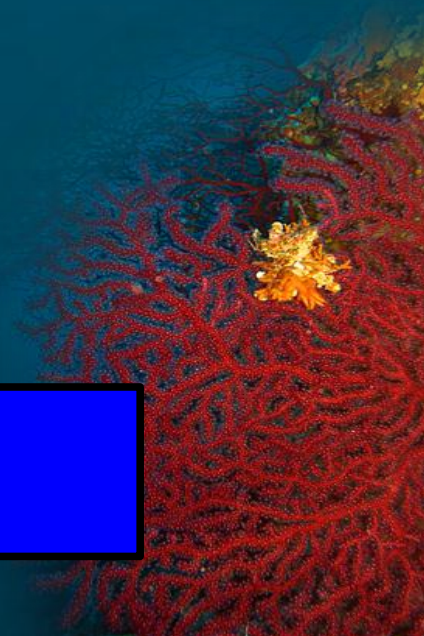
about 15 pixels central part

about 120 pixels halo part

objects present **one or several flashes** of about 0.3 seconds

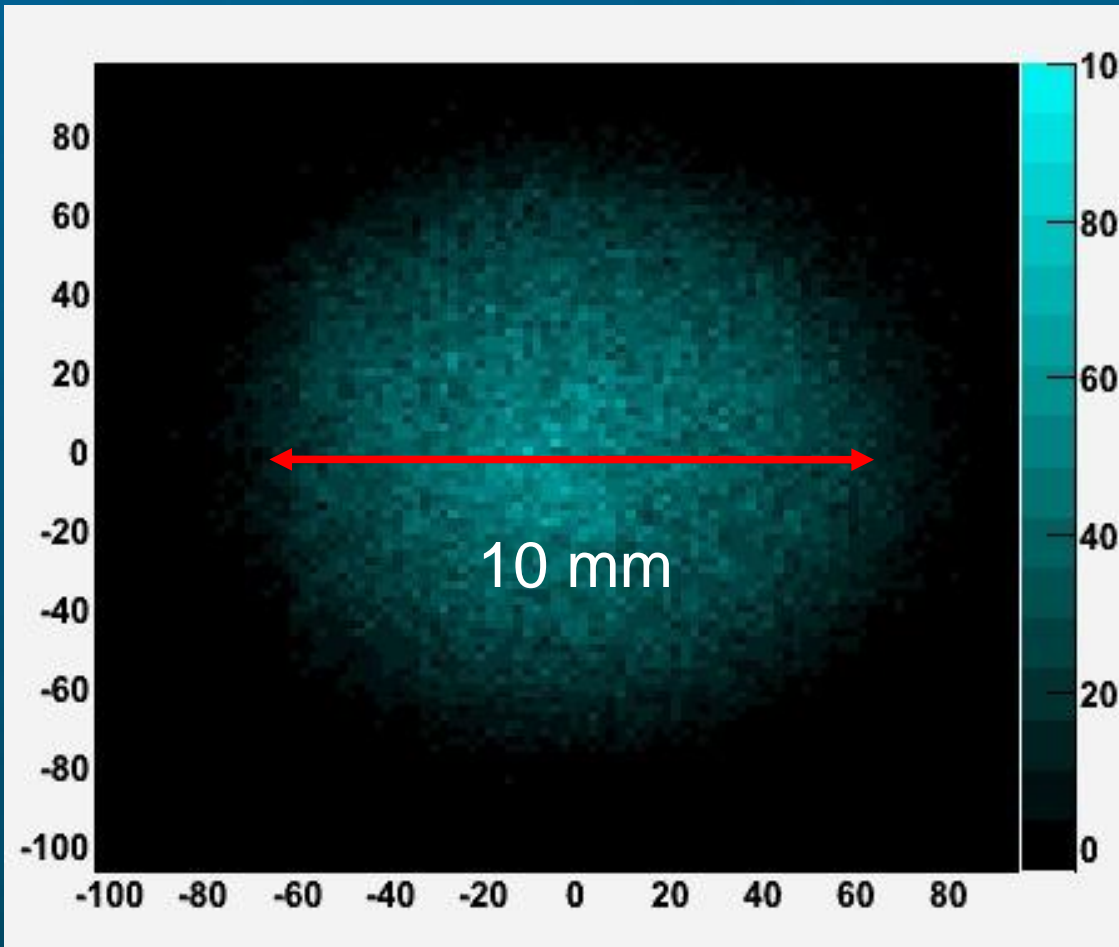
kind of filaments structure

all objects exhibit similar patterns and all objects can be piled up in order to extract a typical size.



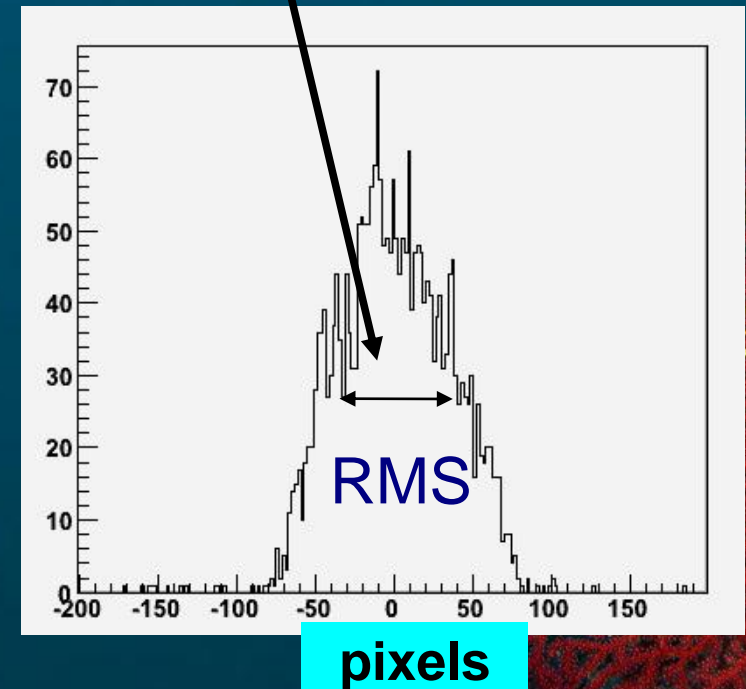


# Object's size on glass window



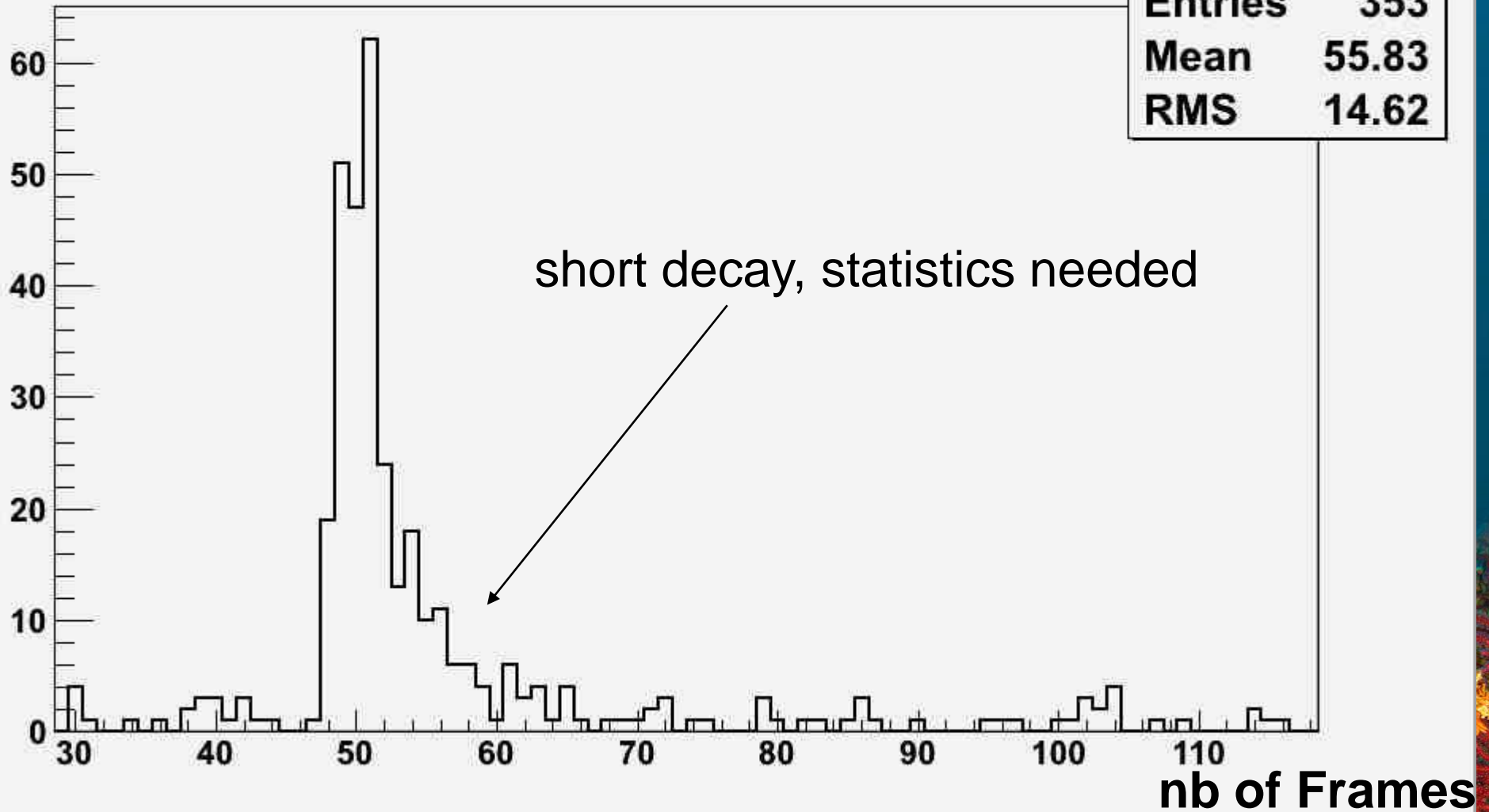
all events statistics, hit positions centered around C.O.G.

- a typical size is extracted
- from the RMS of the distribution :
- **36 pixels.**
- taking into account lens magnification (8) , the real object RMS size is **~ 3 mm**
- max extension of the object : **~ 10 mm**



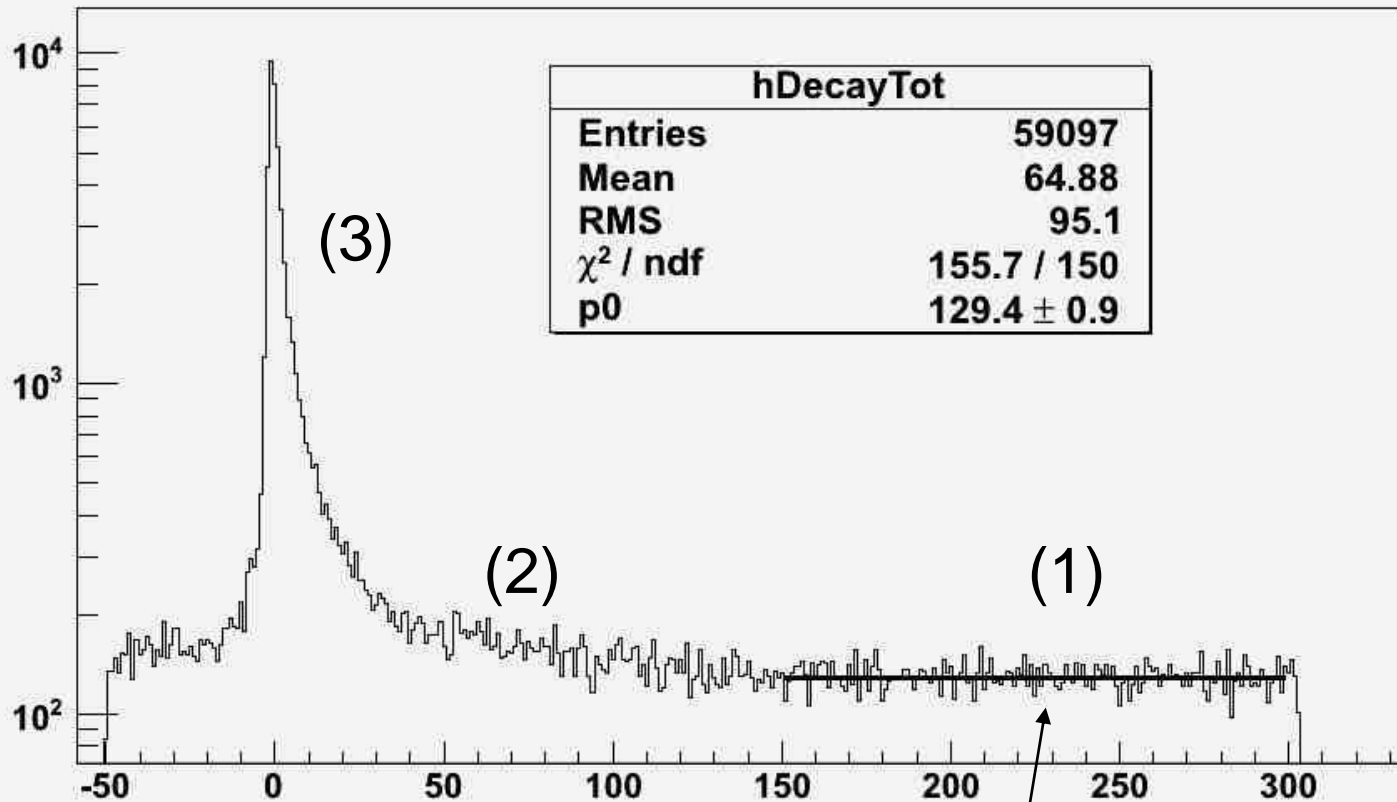
# time dependency

decay



new event selection : Single flash, no motion of the object's C.O.G  
→ 165 events

# Bioluminescence decay (1) : photoelectron noise determination

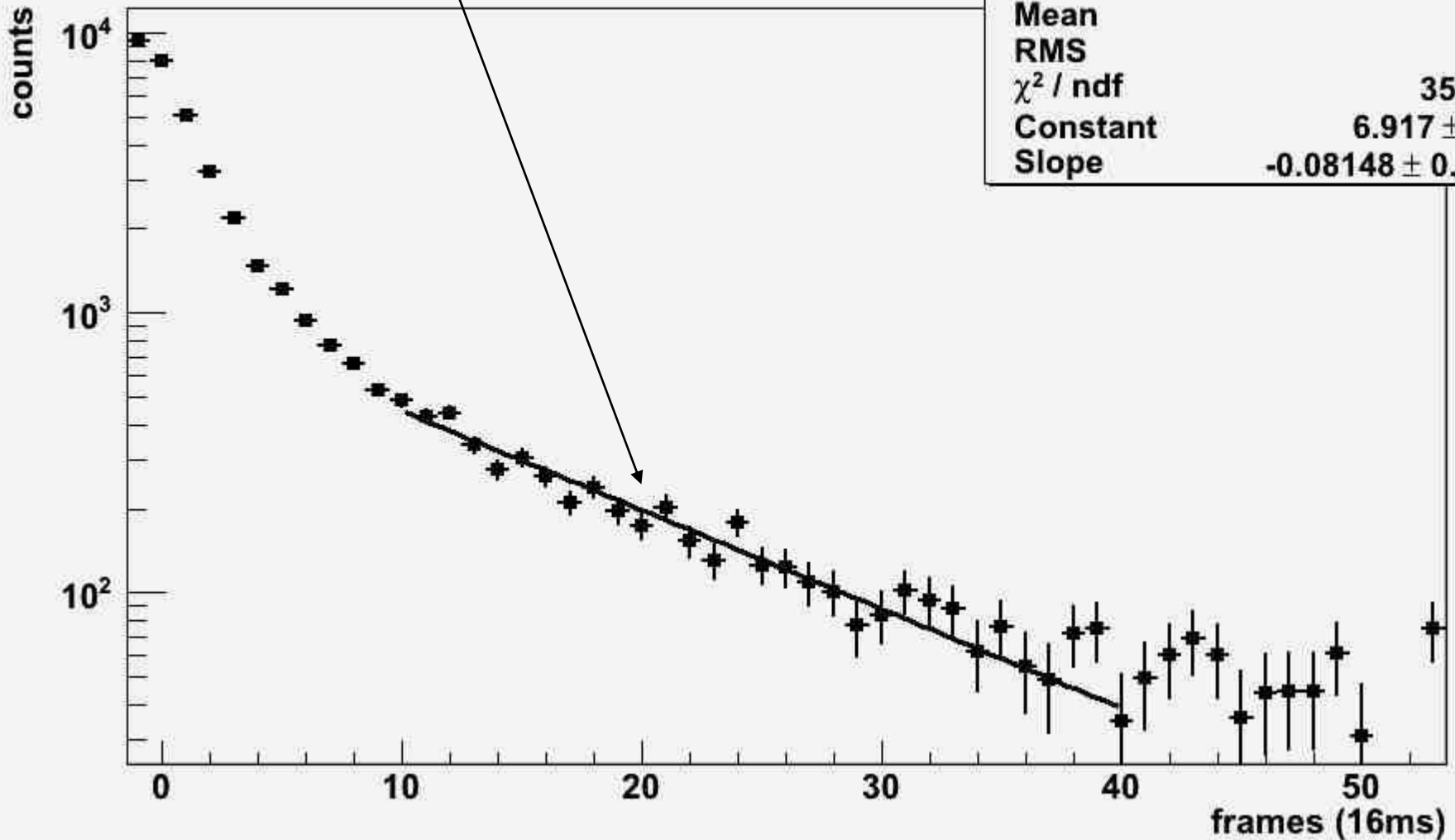


noise : 129 p.e. for 165 events  $\rightarrow$  0.78 p.e. per frame.  
p.e noise is consistent with previous determination  
(slide 13)

# bioluminescence decay

(2) : noise is subtracted, long period part

period = ( 196 +/- 8) ms



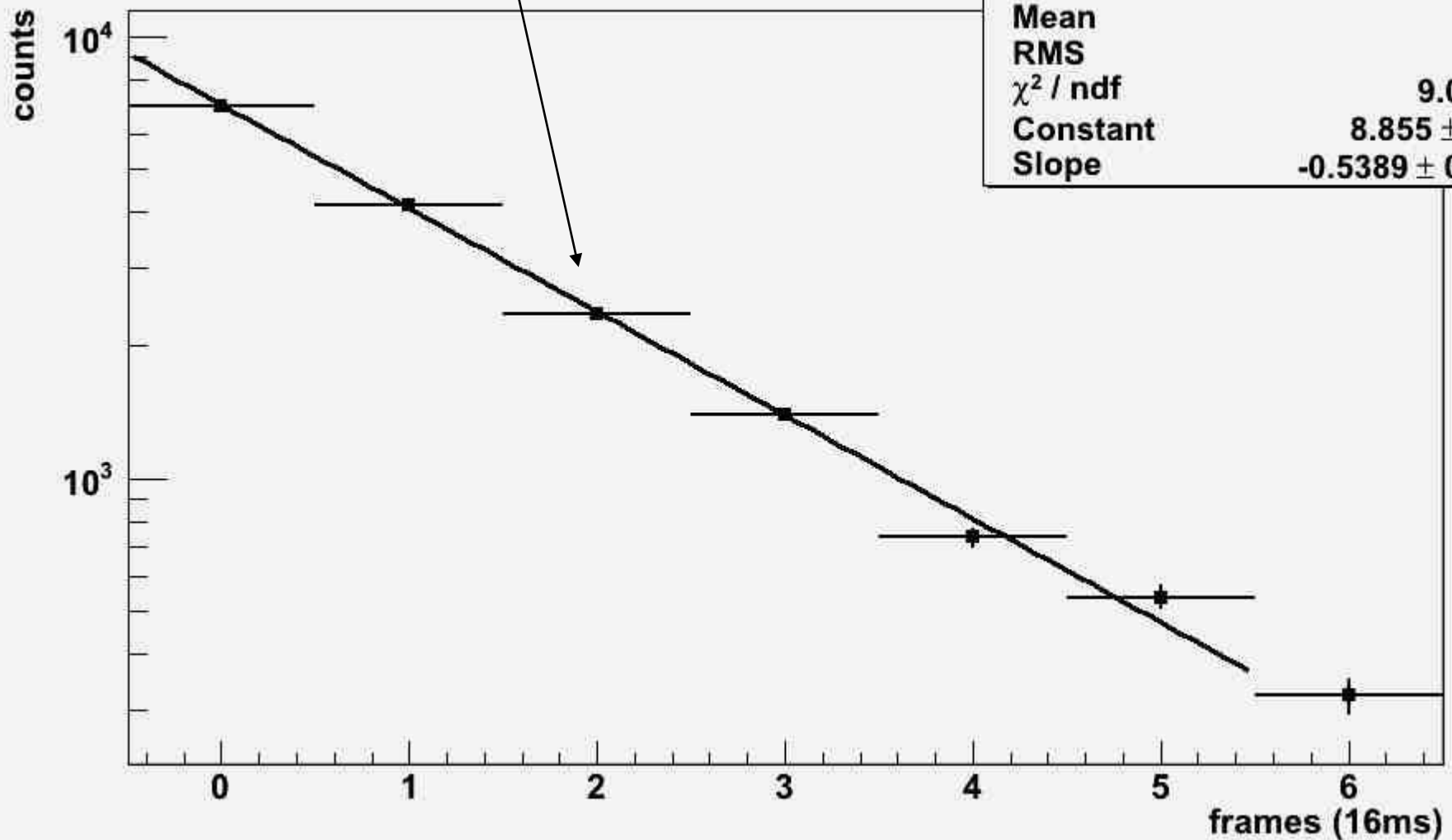
hDecayTot

Entries	59416
Mean	4.399
RMS	8.977
$\chi^2 / \text{ndf}$	35.3 / 29
Constant	$6.917 \pm 0.061$
Slope	$-0.08148 \pm 0.00340$

# Bioluminescence decay

(3) : long period is subtracted, short period part

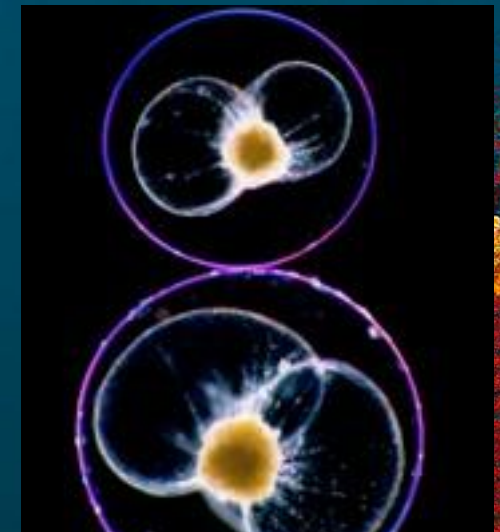
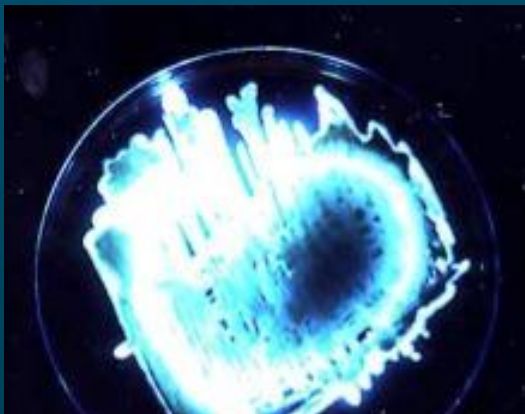
period =  $(29.7 \pm 0.4)$  ms



hDecayTot	
Entries	59735
Mean	1.252
RMS	1.501
$\chi^2 / \text{ndf}$	9.062 / 4
Constant	$8.855 \pm 0.011$
Slope	$-0.5389 \pm 0.0074$

# bioluminescence decay summary

- 165 selected events, 59.000 photoelectrons
  - static events (cog velocity = 0)
  - 1 single flash
- A few mm object's size
- 2 mixed decays are found :
  - $T_{\text{long}} = (196 \pm 8) \text{ ms}$
  - $T_{\text{short}} = (29.7 \pm 0.4) \text{ ms}$

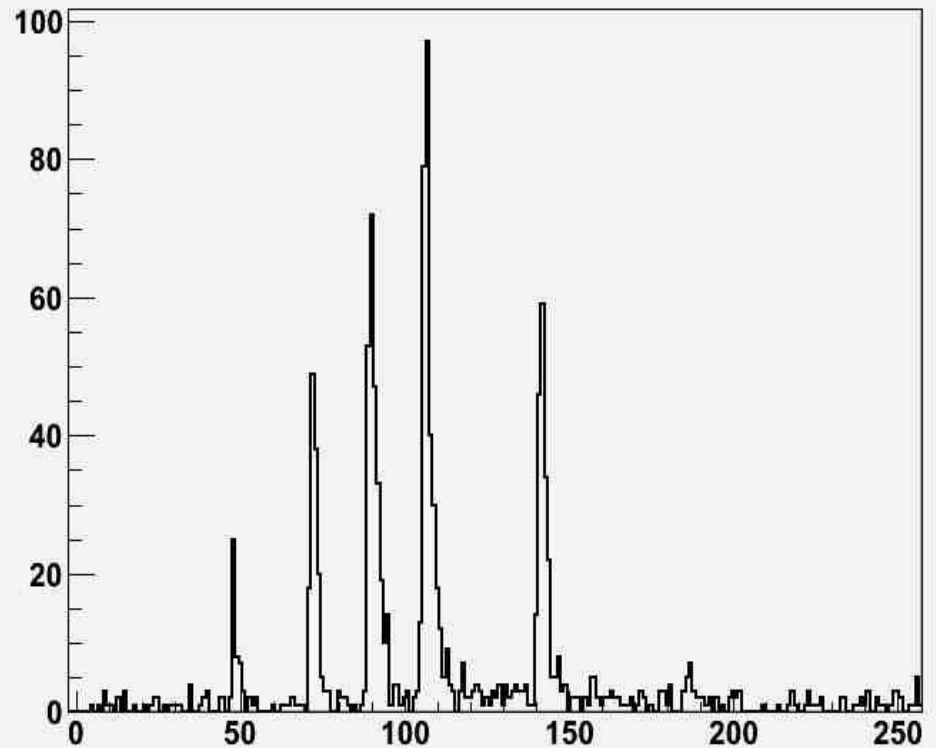
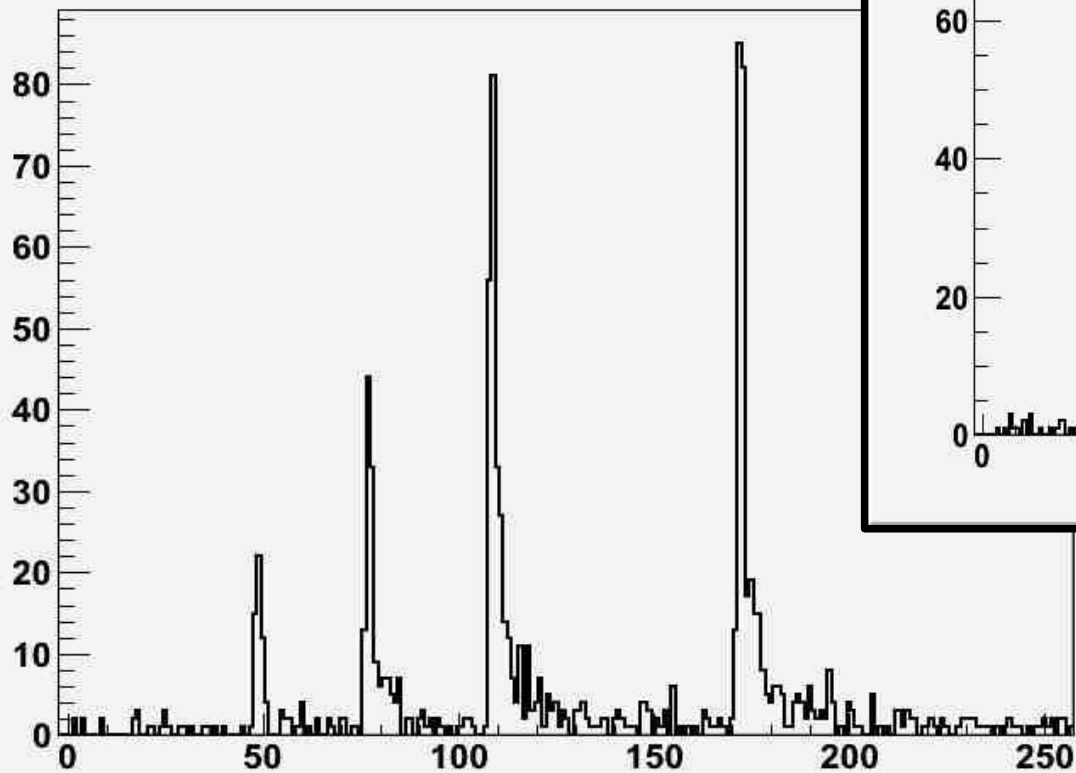


Probably plankton bioluminescence, to be confirmed ..

# flashing events

Analysis in progress ...

photoelectrons number



photoelectrons number

# Conclusion

- cam has been successfully produced in 4 months (instr. electronic, DAQ, computing)
- it's working 24h/24h :
  - only 14 photons are needed to trigger one event
- bioluminescence events are observed and are still under study
  - further studies are in progress to reach a full identification of these objects :
  - optical part has to be improved, are objects well focused or not ?
  - consequences on the real size (critical for the identification)
- correlation between images and sea conditions (current speed and direction, temperature) is in progress thanks to a granted access to ANTARES data (JB) even though it's 500m far from MII and few hundreds meters upper
- Upgrade next year :
  - Lens change to improve field of view
  - Real calibration of sizes, autofocus ?







